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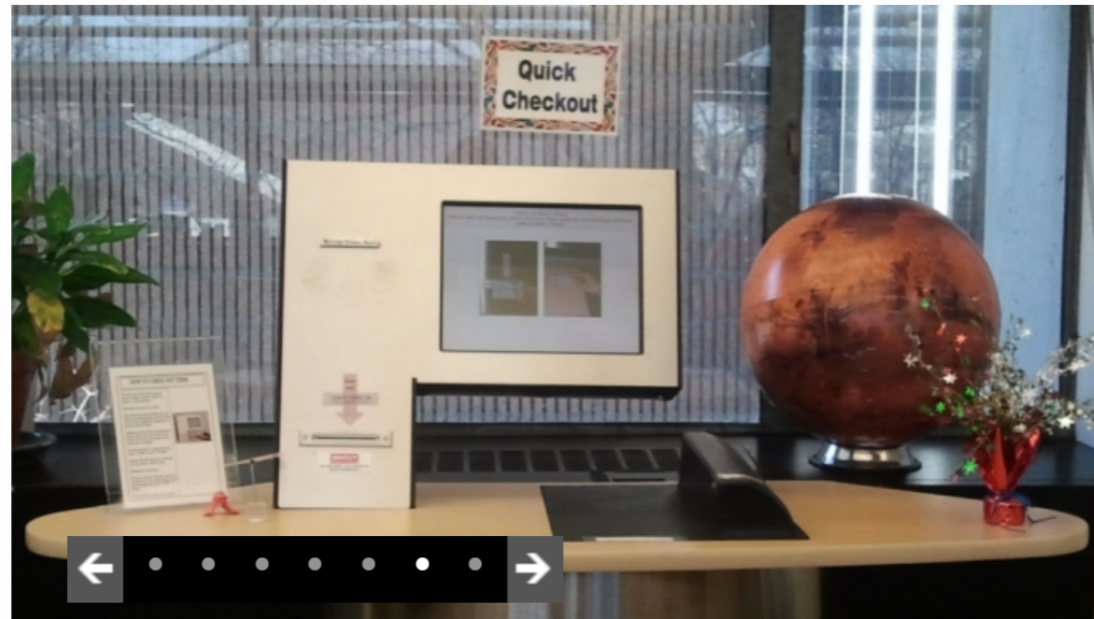
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1610



SIDEREUS NUNCIUS

On the third, at the seventh hour, the sequence. The eastern one was 1 minute, the closest western one 2 minutes; and the

East * ○ * West

to minutes removed from this one. They were absolutely on the same straight line and of equal magnitude.

On the fourth, at the second hour, there were four stars around Jupiter, two to the east and two to the west, and arranged precisely

East * ○ * West

on a straight line, as in the adjoining figure. The easternmost was distant 3 minutes from the next one, while this one was 40 seconds from Jupiter; Jupiter was 4 minutes from the nearest western one, and this one 6 minutes from the westernmost one. Their magnitudes were nearly equal; the one closest to Jupiter appeared a little smaller than the rest. But at the seventh hour the eastern star was 30 seconds apart. Jupiter was 2 minutes from the

East ** ○ **

one, while he was 4 minutes from the next western one was 3 minutes from the westernmost one. They and extended on the same straight line along the ecliptic.

On the fifth, the sky was cloudy.

On the sixth, only two stars appeared flanking Jupiter

East * ○ *

in the adjoining figure. The eastern one was 2 minutes from the next western one 3 minutes from Jupiter. They were on the same straight line with Jupiter and equal in magnitude.

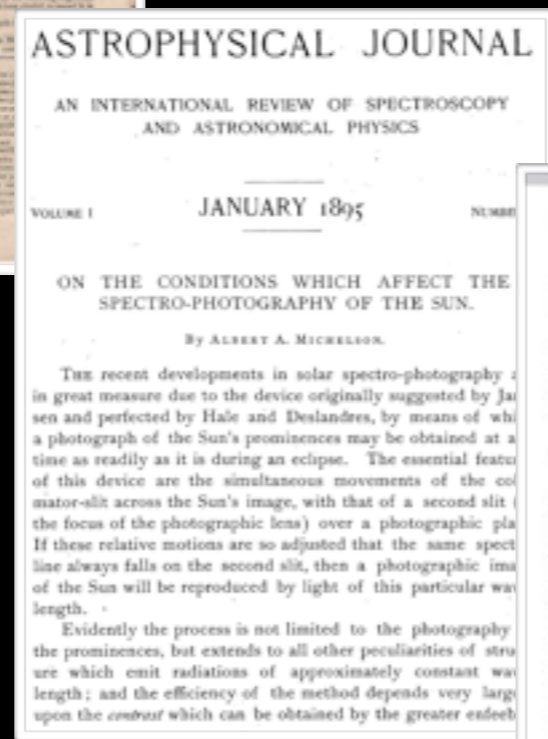
On the seventh, two stars stood near Jupiter, both arranged in this manner.

1665

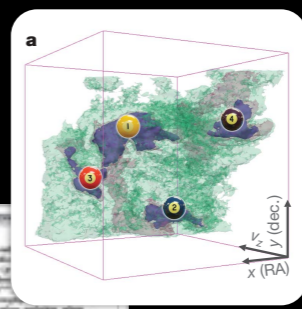
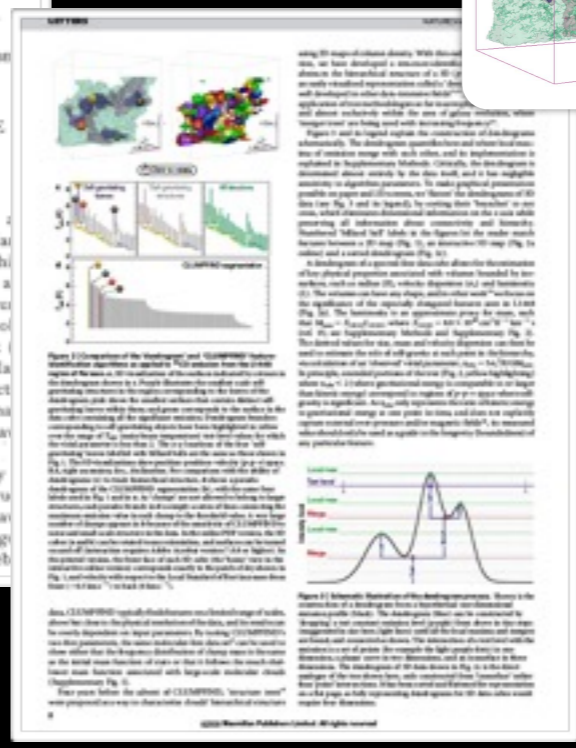


SCHOLARLY COMMUNICATION

1895



2009



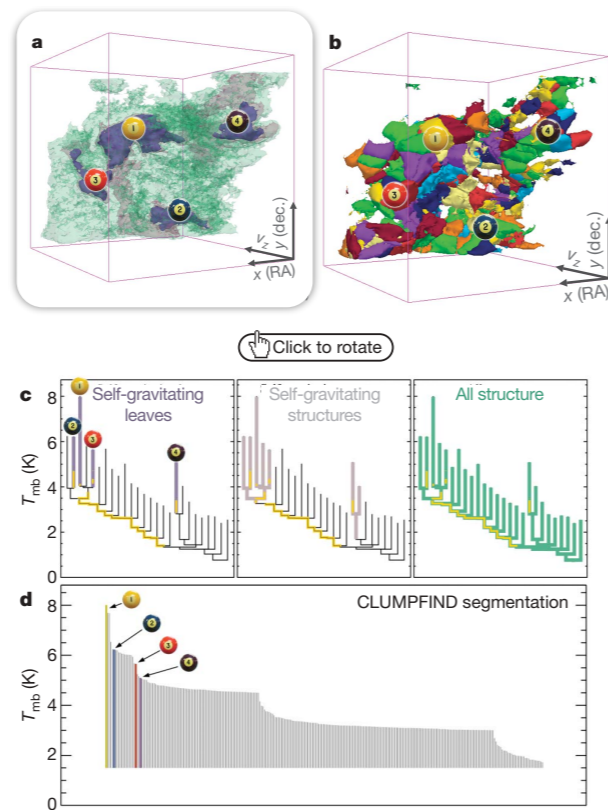


Figure 2 | Comparison of the 'dendrogram' and 'CLUMPFIND' feature-identification algorithms as applied to ^{13}CO emission from the L1448 region of Perseus. **a**, 3D visualization of the surfaces indicated by colours in the dendrogram shown in **c**. Purple illustrates the smallest scale self-gravitating structures in the region corresponding to the leaves of the dendrogram; pink shows the smallest surfaces that contain distinct self-gravitating leaves within them; and green corresponds to the surface in the data cube containing all the significant emission. Dendrogram branches corresponding to self-gravitating objects have been highlighted in yellow over the range of T_{mb} (main-beam temperature) test-level values for which the virial parameter is less than 2. The x - y locations of the four 'self-gravitating' leaves labelled with billiard balls are the same as those shown in Fig. 1. The 3D visualizations show position-position-velocity (p - p - v) space. RA, right ascension; dec., declination. For comparison with the ability of dendrograms (**c**) to track hierarchical structure, **d** shows a pseudo-dendrogram of the CLUMPFIND segmentation (**b**), with the same four labels used in Fig. 1 and in **a**. As 'clumps' are not allowed to belong to larger structures, each pseudo-branch in **d** is simply a series of lines connecting the maximum emission value in each clump to the threshold value. A very large number of clumps appears in **b** because of the sensitivity of CLUMPFIND to noise and small-scale structure in the data. In the online PDF version, the 3D cubes (**a** and **b**) can be rotated to any orientation, and surfaces can be turned on and off (interaction requires Adobe Acrobat version 7.0.8 or higher). In the printed version, the front face of each 3D cube (the 'home' view in the interactive online version) corresponds exactly to the patch of sky shown in Fig. 1, and velocity with respect to the Local Standard of Rest increases from front (-0.5 km s^{-1}) to back (8 km s^{-1}).

data, CLUMPFIND typically finds features on a limited range of scales, above but close to the physical resolution of the data, and its results can be overly dependent on input parameters. By tuning CLUMPFIND's two free parameters, the same molecular-line data set⁸ can be used to show either that the frequency distribution of clump mass is the same as the initial mass function of stars or that it follows the much shallower mass function associated with large-scale molecular clouds (Supplementary Fig. 1).

Four years before the advent of CLUMPFIND, 'structure trees'⁹ were proposed as a way to characterize clouds' hierarchical structure

using 2D maps of column density. With the help of 2D work as inspiration, we have developed a structure-identification algorithm that abstracts the hierarchical structure of a data cube into a form that is an easily visualized representation called a 'dendrogram'. This method is well developed in other data-intensive applications, such as image segmentation, and almost exclusively within the area of computer vision. 'merger trees' are being used with increasing frequency.

Figure 3 and its legend explain the dendrogram process schematically. The dendrogram of a data cube is determined almost entirely by the choice of a test level, and its sensitivity to algorithm parameters is limited. The dendrogram is possible on paper and 2D screen, and its construction is a simple process, which eliminates dimensions, while preserving all information. The dendrogram is a 'billiard ball' labelled with features between a 2D map (online) and a sorted dendrogram.

A dendrogram of a spectral cube is a tree diagram of key physical properties, such as radius (R), velocity dispersion (σ_v), and luminosity (L). The volumes can have any shape, and the significance of the especially elongated features is determined by their luminosity (Fig. 2a). The luminosity is an approximate proxy for mass, such that $M_{\text{lum}} = X_{13\text{CO}} L_{13\text{CO}}$, where $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^{-2} \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter, $\alpha_{\text{obs}} = 5\sigma_v^2 R / GM_{\text{lum}}$. In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{\text{obs}} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p - p - v space where self-gravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields¹⁶, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.

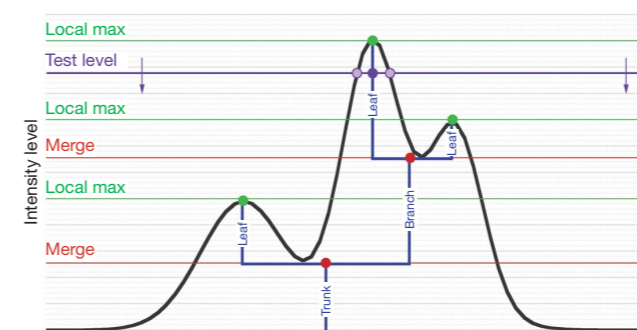


Figure 3 | Schematic illustration of the dendrogram process. Shown is the construction of a dendrogram from a hypothetical one-dimensional emission profile (black). The dendrogram (blue) can be constructed by 'dropping' a test constant emission level (purple) from above in tiny steps (exaggerated in size here, light lines) until all the local maxima and mergers are found, and connected as shown. The intersection of a test level with the emission is a set of points (for example the light purple dots) in one dimension, a planar curve in two dimensions, and an isosurface in three dimensions. The dendrogram of 3D data shown in Fig. 2c is the direct analogue of the tree shown here, only constructed from 'isosurface' rather than 'point' intersections. It has been sorted and flattened for representation on a flat page, as fully representing dendrograms for 3D data cubes would require four dimensions.

2009

3D PDF

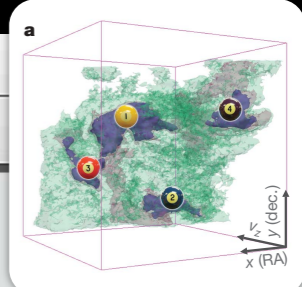
INTERACTIVITY IN
A "PAPER"

A role for self-gravity at multiple length scales in the process of star formation

Alyssa A. Goodman^{1,2}, Erik W. Rosolowsky^{2,3}, Michelle A. Borkin^{1,4}, Jonathan B. Foster², Michael Halle^{1,4}, Jens Kauffmann^{1,2} & Jaime E. Pineda²

Self-gravity plays a decisive role in the final stages of star formation, where dense cores (size ~ 0.1 parsecs) inside molecular clouds collapse to form star-plus-disk systems. But self-gravity's role at earlier times (and on larger length scales, such as ~ 1 parsecs) is unclear: some molecular cloud simulations that do not include self-gravity suggest that 'turbulent fragmentation' alone is sufficient to create a mass distribution of dense cores that resembles, and sets, the stellar initial mass function. In particular, more self-gravity plays a significant role over the full range of possible scales traced by ^{13}CO observations in the L1448 molecular cloud, but not everywhere in the observed region. In particular, more than 90 per cent of the compact 'pre-stellar cores' traced by peaks of dust emission are projected on the sky within one of the dendrogram's self-gravitating 'leaves'. As these peaks mark the locations of key physical properties, such as radius (R), velocity dispersion (σ_v), and luminosity (L), the significance of the especially elongated features is determined by their luminosity (Fig. 2a). The luminosity is an approximate proxy for mass, such that $M_{\text{lum}} = X_{13\text{CO}} L_{13\text{CO}}$, where $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^{-2} \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter, $\alpha_{\text{obs}} = 5\sigma_v^2 R / GM_{\text{lum}}$. In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{\text{obs}} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p - p - v space where self-gravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields¹⁶, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.

Goodman et al. 2009, Nature,
cf. Fluke et al. 2009



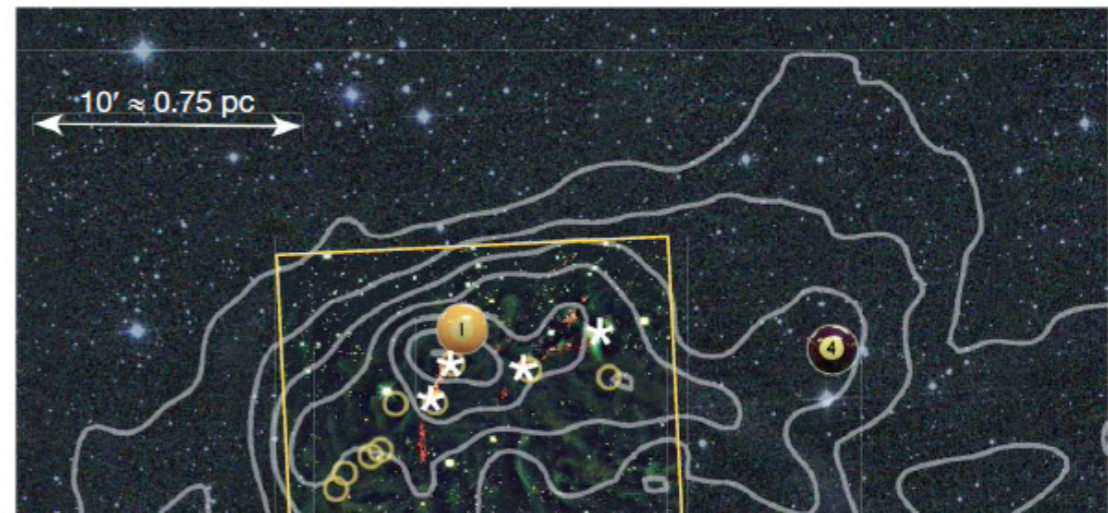
LETTERS

A role for self-gravity at multiple length scales in the process of star formation

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overlapping features as an option, significant emission found between prominent clumps is typically either appended to the nearest clump or turned into a small, usually 'pathological', feature needed to encompass all the emission being modelled. When applied to molecular-line



1610



SCHOLARLY COMMUNICATION



SIDEREUS NUNCIUS

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East * ○ * * West

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East * ○ *

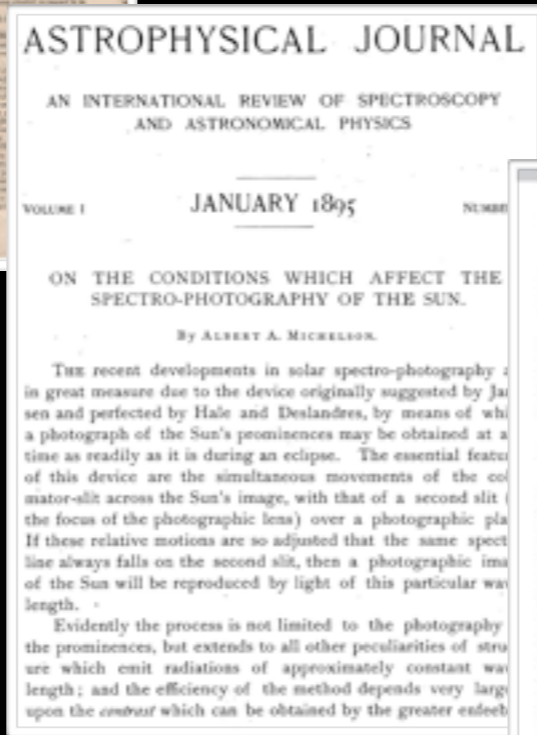
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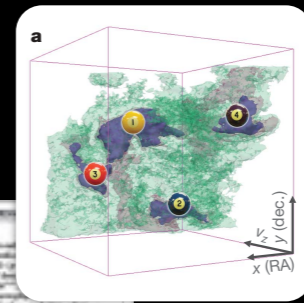
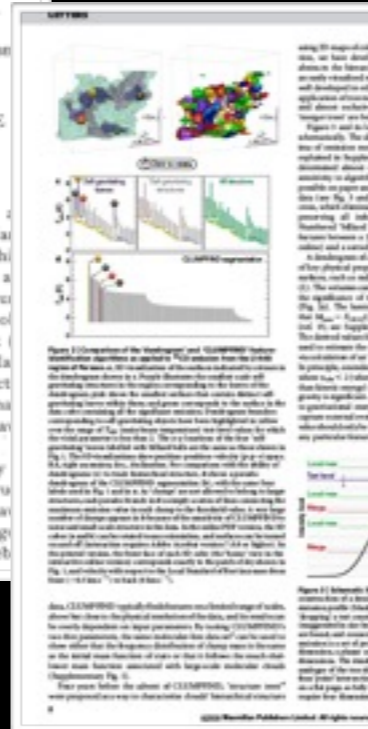
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1895



2009



2015

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The "Paper" of the Future

Alyssa Goodman, Josh Peek, Alberto Accomazzi, Chris Beaumont, Christine L. Borgman, How-Huan Hope Chen, Merce Crossas, Christopher Erdmann, August Muench, Alberto Pepe, Curtis Wong

A 5-minute video demonstration of this paper is available at this YouTube link.

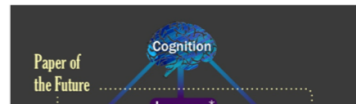
1 Preamble

A variety of research on human cognition demonstrates that humans learn and communicate best when more than one processing system (e.g. visual, auditory, touch) is used. And, related research also shows that, no matter how technical the material, most humans also retain and process information best when they can put a narrative "story" to it. So, when considering the future of scholarly communication, we should be careful not to do bithely away with the linear narrative format that articles and books have followed for centuries: instead, we should enrich it.

Much more than text is used to communicate in Science. Figures, which include images, diagrams, graphs, charts, and more, have enriched scholarly articles since the time of Galileo, and ever-growing volumes of data underpin most scientific papers. When scientists communicate face-to-face, as in talks or small discussions, these figures are often the focus of the conversation. In the best discussions, scientists have the ability to manipulate the figures, and to access underlying data. In real-time, so as to test out various what-if scenarios, and to explain findings more clearly. **This short article explains—and shows with demonstrations—how scholarly "papers" can morph into long-lasting rich records of scientific discourse,** enriched with deep data and code linkages, interactive figures, audio, video, and commenting.

Comments:

- Konrad Hinsen 3 days ago - Public: Many good suggestions, but if the goal is "long-lasting rich records of scientific discourse", a more careful and critical attitude towards electronic artifacts is appropriate. I do see it concerning videos, but not so worried on the much more critical situation in software. Archiving source code is not sufficient: all the dependencies, plus the complete build environment, would have to be conserved as well to make things work a few years from now. An "executable figure" in the form of an IPython notebook will...
- Merce Crossas 3 days ago - Public: Konrad, good points, this has been a concern for the community working on reproducibility. Regarding data repositories, Dataverse handles long-term preservation and access of data files in the following way: 1) for some data files that the repository recognizes (such as R Data, SPSS, STATA), which depend on a statistical package, the system converts them into a preservation format (such as a tab/CSV format). Even though the original format is also saved and can be accessed, the new preservation format gu...
- Konrad Hinsen 1 day ago - Public: That sounds good. I hope more repositories will follow the example of Dataverse. Figure in particular has a very different attitude, encouraging researchers to deposit as much as possible. That's perhaps a good strategy to change habits, but in the long run it could well backfire when people find out in a few years that 80% of those deposits have become useless.
- Christine L. Borgman 4 months ago - Private: "Substitutions"



The "Paper" of the Future

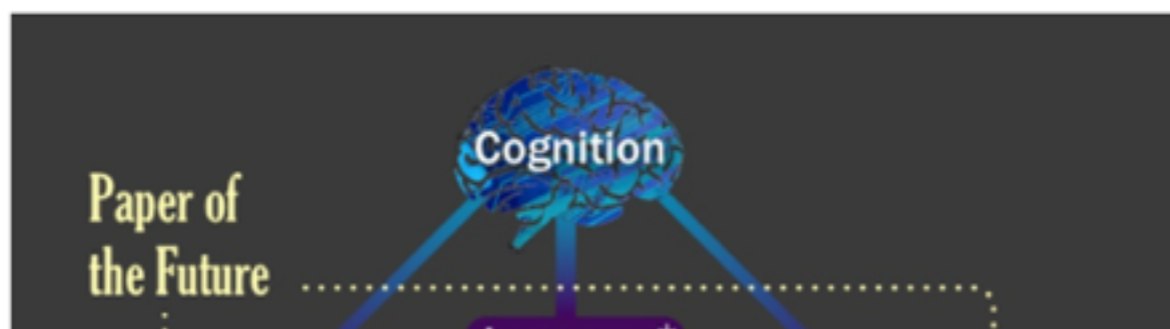
Alyssa Goodman, Josh Peek, Alberto Accomazzi, Chris Beaumont, Christine L. Borgman, How-Huan Hope Chen, Merce Crosas, Christopher Erdmann, August Muench, Alberto Pepe, Curtis Wong [+ Add author](#) [✕ Re-arrange authors](#)

A 5-minute video demonstration of this paper is available at [this YouTube link](#).

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3

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2

Merce Crosas 3 days ago · Public

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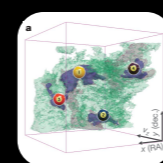
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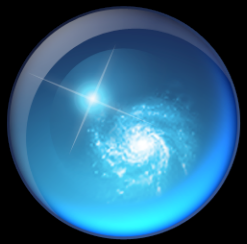
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Christine L. Borgman 4 months ago · Private

"publications"





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The screenshot shows the WorldWide Telescope interface with several key components:

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- Collections:** A row of thumbnails for 'All-Sky Surveys' including 'Digitized Sky Survey', 'VLSS: VLA Low-frequency Sky Survey', 'WMAP ILC 5-Year', 'SFD Dust Map (Infrared)', 'IRIS: Improved Resolution', '2MASS: Two Micron All Sky Survey', and 'Hydrogen Alpha Full Sky Survey'.
- Finder Scope:** A central circular view showing a galaxy with a crosshair. Below it, a panel displays details for 'NGC224', including its classification as a 'Spiral Galaxy in Andromeda' and various astronomical coordinates (RA, Dec, Alt, Az, Rise, Transit, Set).
- Context Bar:** Located at the bottom, it shows 'Look At' (set to 'Sky'), 'Imagery' (set to 'Digitized Sky Survey'), and a 'Context globe' showing the current field of view in the Andromeda constellation.
- Image Credits:** A panel at the bottom provides information about the data source: 'Data provided by two NASA satellites, the Infrared Astronomy Satellite (IRAS) and the Cosmic Background Explorer (COBE). Processing http://astro.berkeley.edu/~marc/dust/'.

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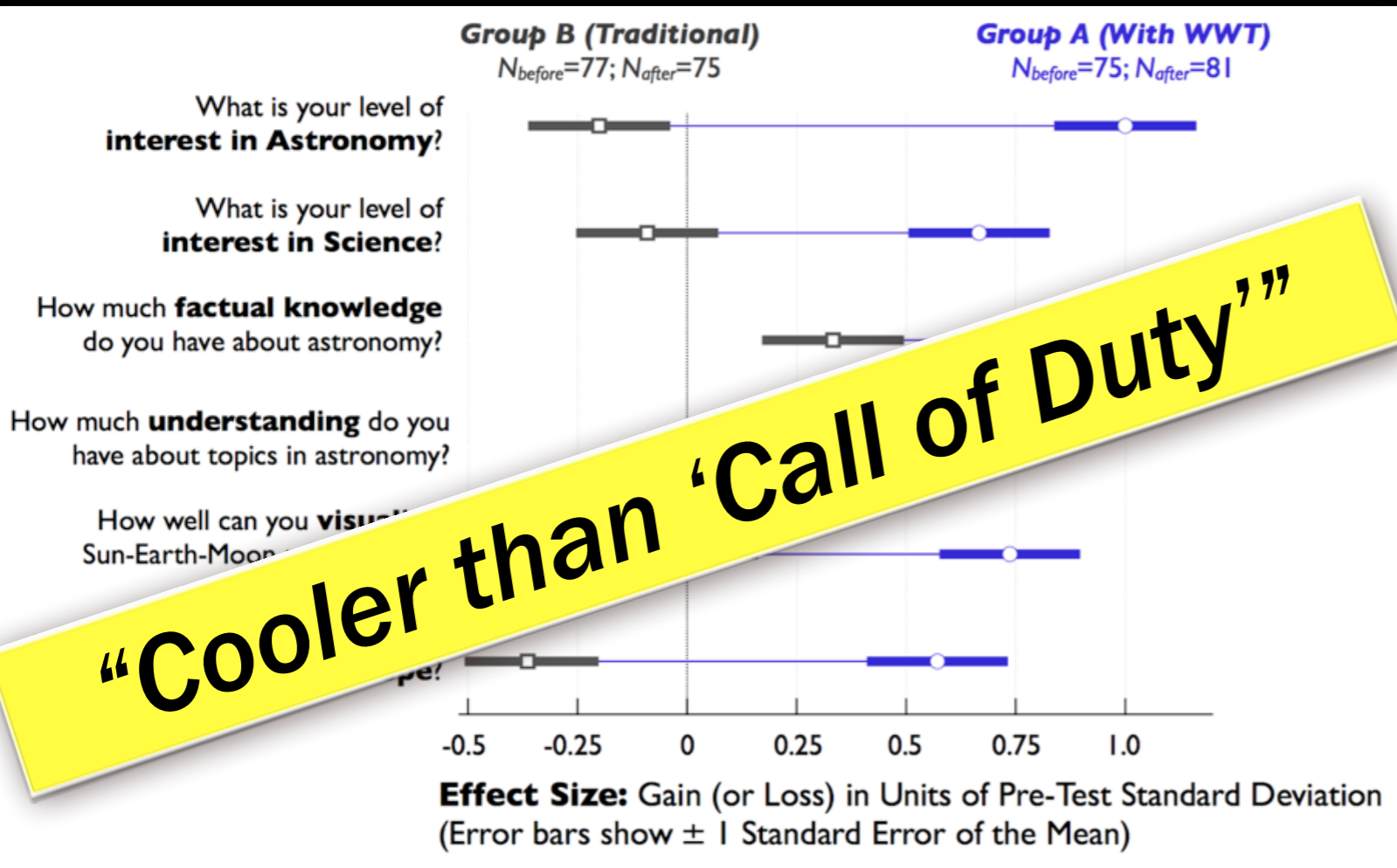
Much more than "just" the sky at night! 3D features can take you to other planets, stars & galaxies.

Finder Scope links to Wikipedia, publications, and data, so you can learn more

Context bar shows items of interest in current field of view

Context globe shows where you're looking.

WWT Ambassadors



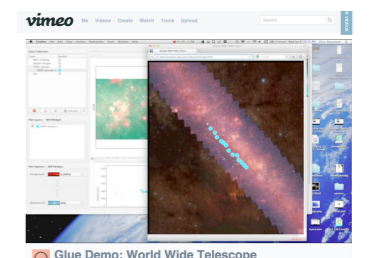
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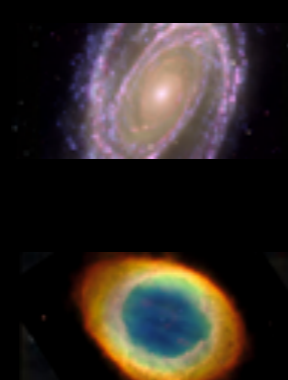
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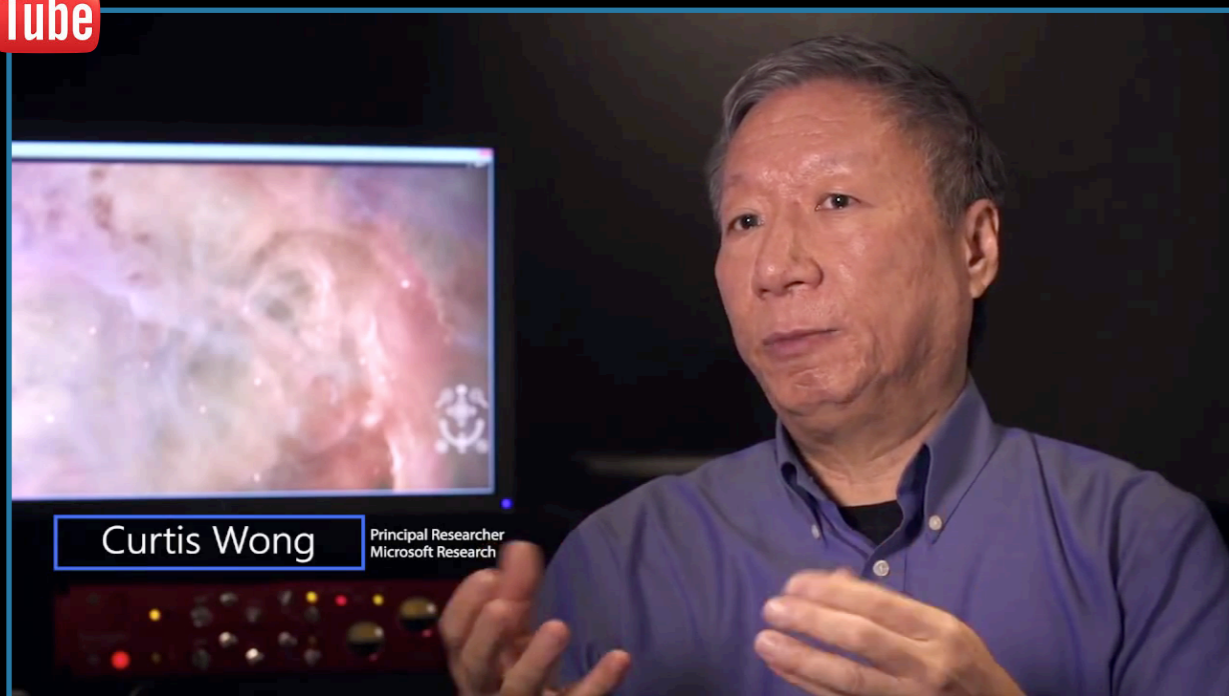
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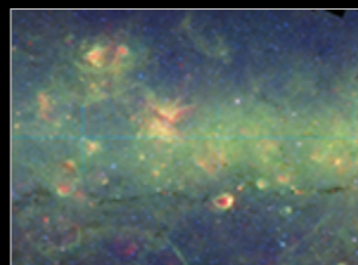


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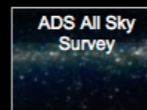
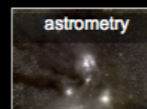
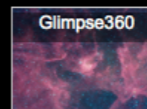
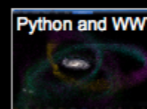
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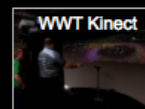


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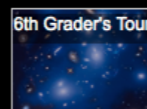


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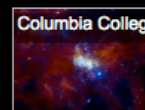
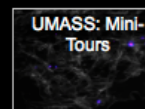
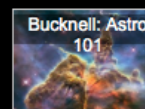


UNIVERSITY



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LEARNING MODULES

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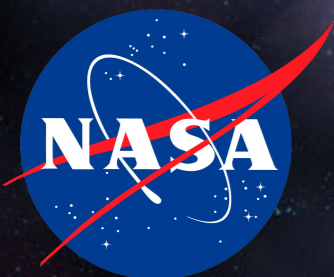


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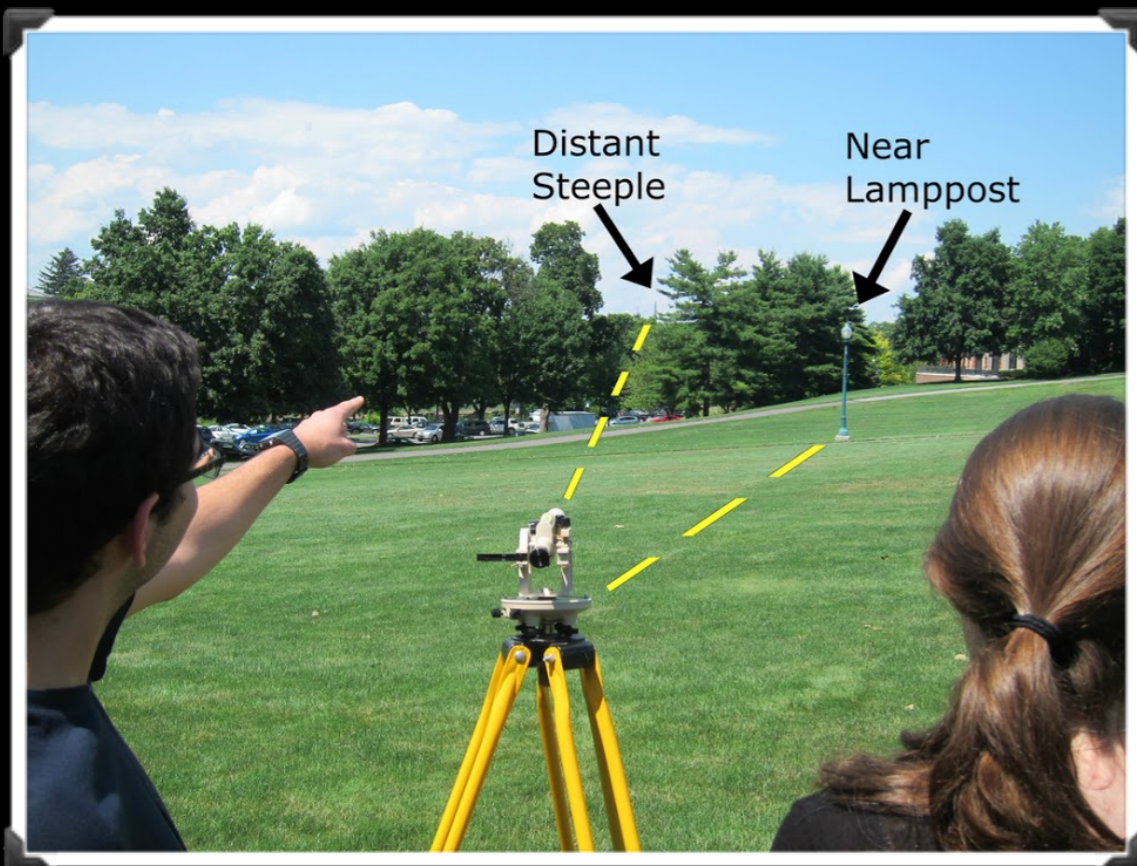


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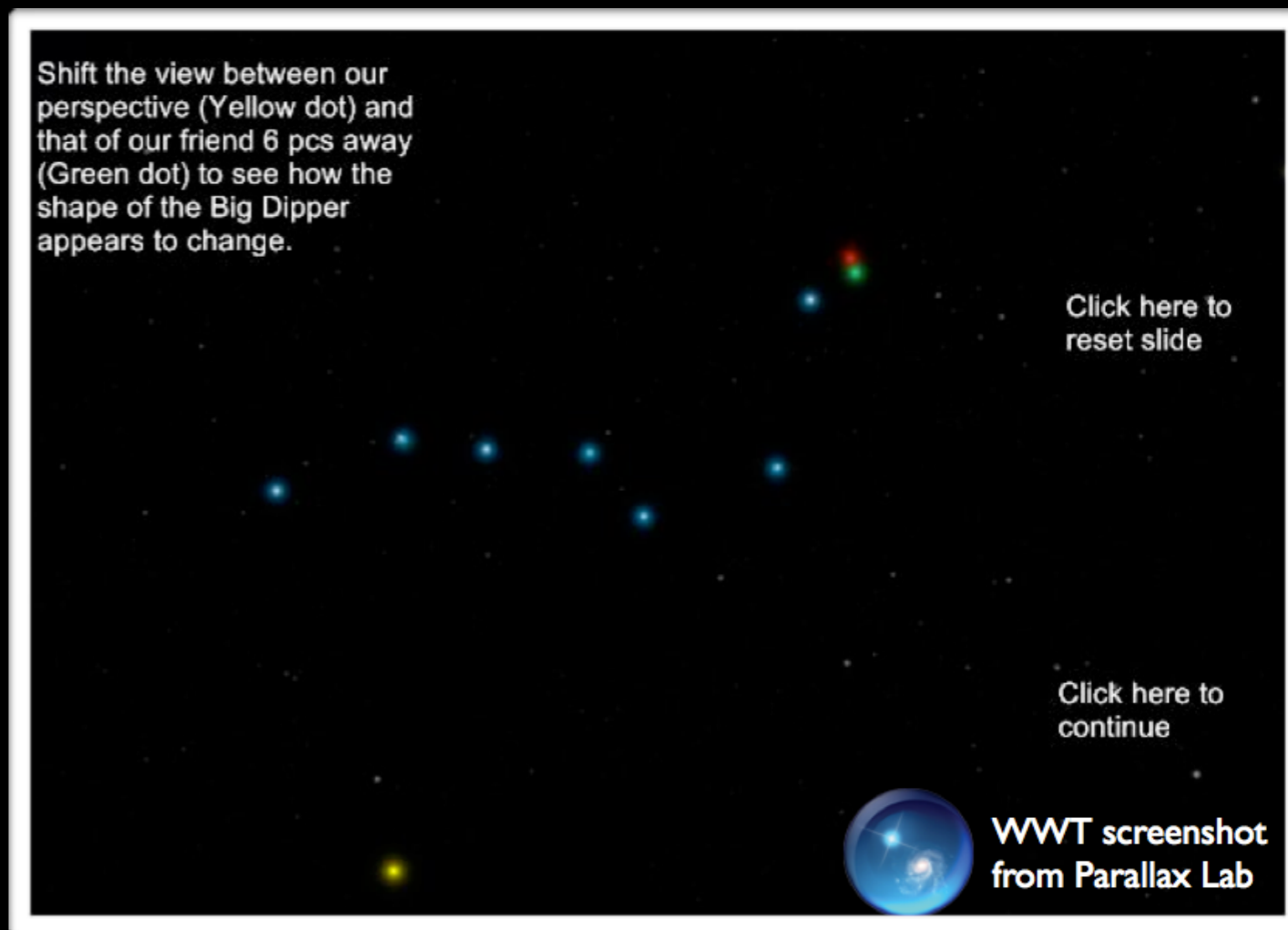
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John
Templeton
Foundation




Shift the view between our perspective (Yellow dot) and that of our friend 6 pcs away (Green dot) to see how the shape of the Big Dipper appears to change.



GRAD STUDENT LEARNING MODULES

- ORIGINS
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Landmarks of the ISM

The Milky Way

3:26

ir Medium

0:12 / 8:24

YouTube

The course structure is such a success that the Harvard Astronomy Department adopts a requirement of all graduate students to create an online outreach project to explain their PhD Thesis work to general audiences. Creating a WorldWide Telescope tours to describe research is a great way students can share their projects with large audiences for years to come.



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- EARTHQUAKES REVEALED

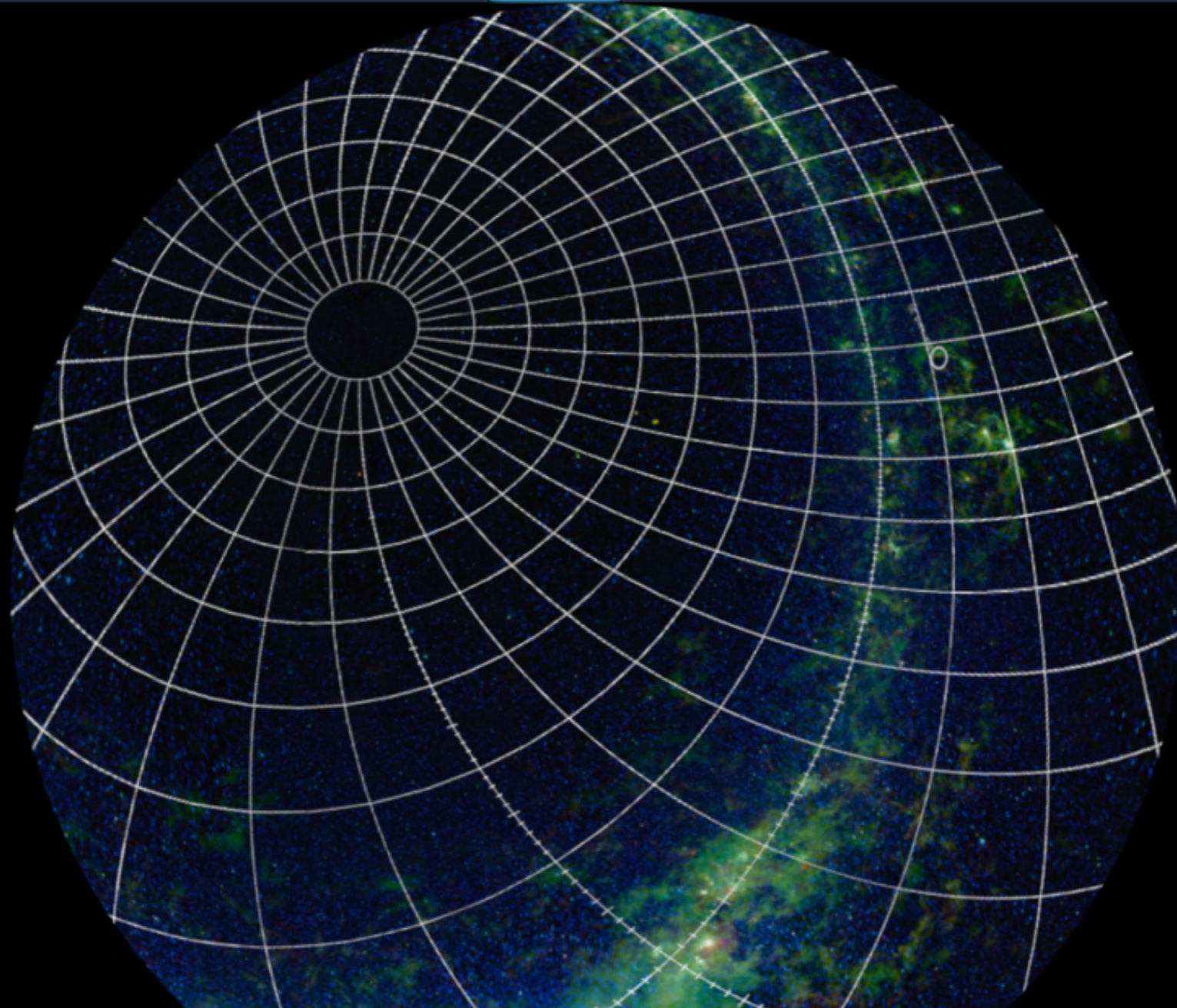
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Look At: Sky | Imagery: WISE All Sky (Infrared) | Context Search Filter: All | 1 of 33

Orion Nebula | Orion Nebula | Orion Nebula | Barnard 30 | **Barnard 30** | Barnard 30 520 | Abell 520 | Messier 42 | Hubble Probes th... | Orion Nebula (N...

RA: 09h36m 16s | Dec: -61:36:28 | Carina 22:40:03



WorldWide Telescope: Planetariums



COSMIC WONDER



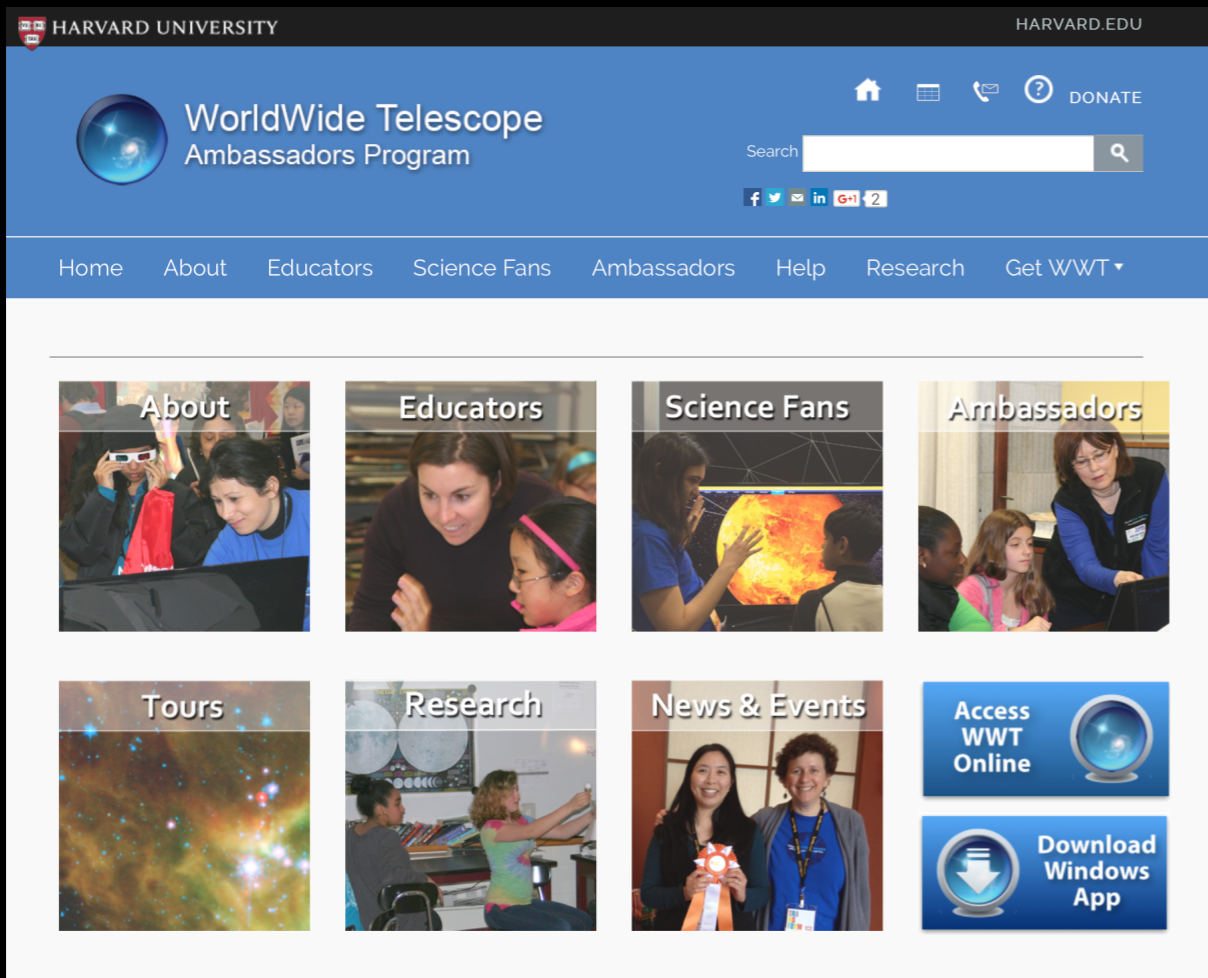
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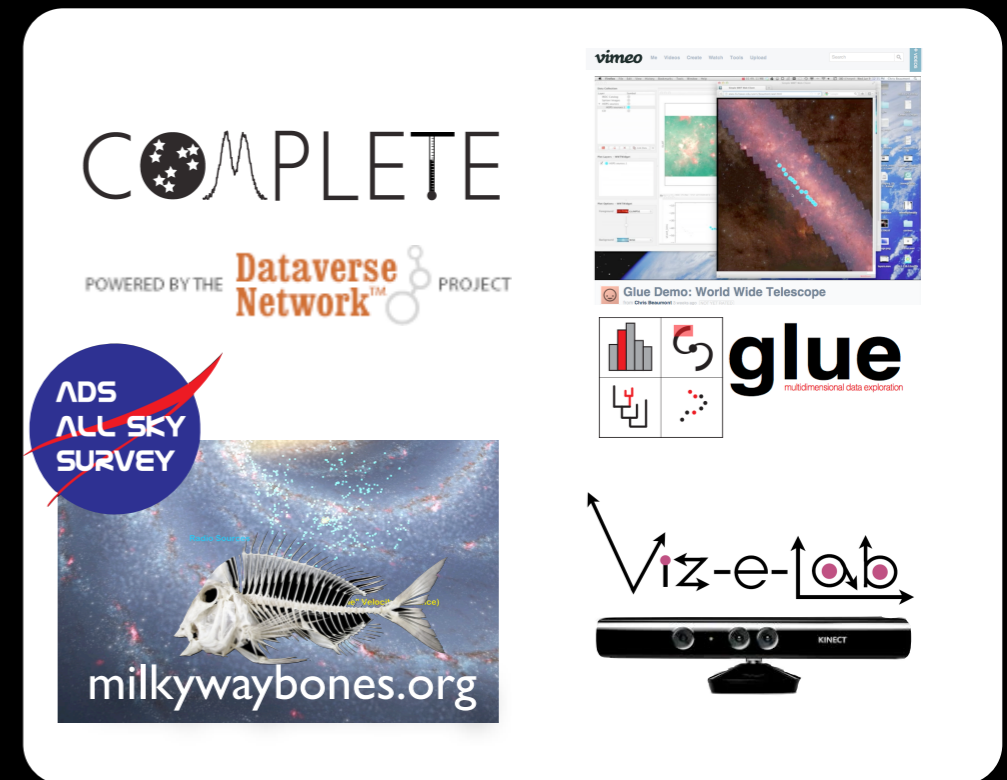
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"Tours"



The Eagle Nebula





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Region: In Perseus and Taurus

ads
NASA

α (2000) 3h 38m 14s, δ (2000) +31° 25'

α (1875) 3h 30m 30s, δ (1875) +31° 00'

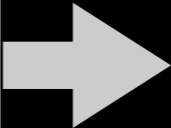
Area
In Perseus and Taurus

Galactic Coordinates
127°, -18°

Scale
1 cm = 18'.2 or 1 in = 46'.2

Chart Plate & Chart
Table Text

enlarge [+] printable PDF



Bar-p1-0003_sm
Bernard's Image of Perseus,
www.library.gatech.edu/spd...

December 11

1

astrometry.net
Hello, this is the blind astrometry solver.
Your results are: (RA, Dec) center:
(54.3006782184, 31.43266374) degrees
Orientation: 5.2134989764 deg E of N
Pixel scale: 18.56371997 arcseconds
Your field contains: NGC 1465 IC 1985 C
Per / Ask o Per 42Per 42Per NGC 1333
IC 348 IC 2003 View in World Wide
Telescope: — If you would like to have
other images solved, please submit
them to the astrometry group.

astrometry (group)



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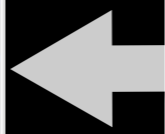
ADS ALL SKY SURVEY

Look At Imagery Info Image Coordinates

Digitial Sky Survey (Color)

California Nebula IC348-IC 348 IC1911 IC1881 IC1985 IC1874 IC1900

Perseus 11/22/24



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Collections > Open Collections > Bar-p1-0003_sm X California Nebula

Bar-p1-0003_sm

Look At Imagery Info Image Coordinates

Digitial Sky Survey (Color)

California Nebula IC348-IC 348 IC1911 IC1881 IC1985 IC1874 IC1900

Perseus 11/22/24



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ophiuchus

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Images

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- The Astrophysical Journal 1,049
- The Astrophysical Journal Letters 55
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Figure sets (3)

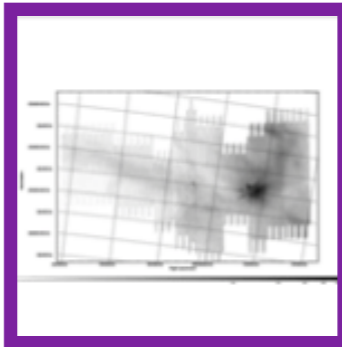
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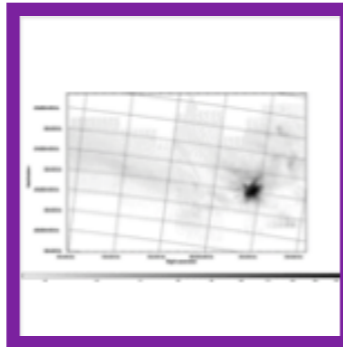
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SHOW: 25

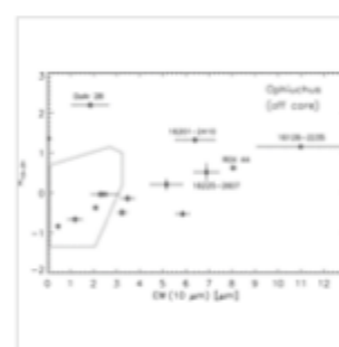
1 of 67



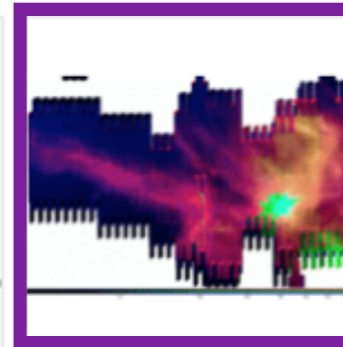
Spitzer MIPS 70 μm mosaic of **Ophiuchus**



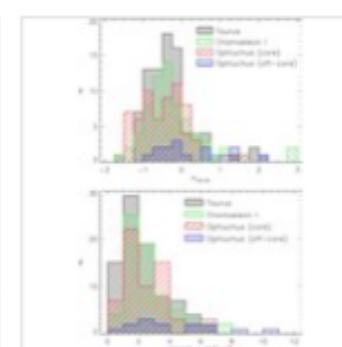
Spitzer MIPS 24 μm mosaic of **Ophiuchus**



Same as Figure 10, but for the **Ophiuchus** off-co...



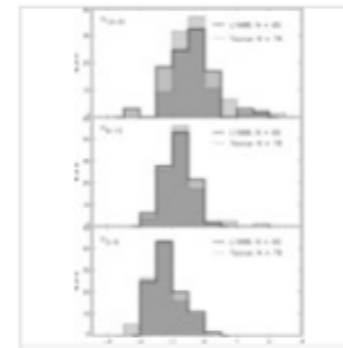
A 24 (blue), 70 (green), and 160 μm (red) mosai...



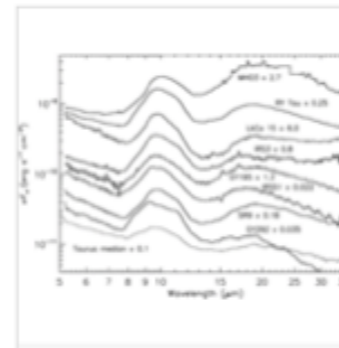
Histograms for the distribution of n13-31 and



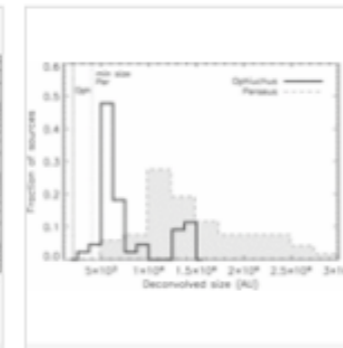
SEDs of WTTS disks in **Ophiuchus**.



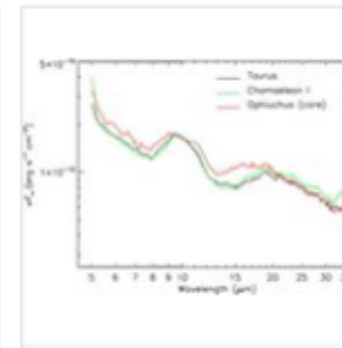
Distribution of n13-31 values for samples in th...



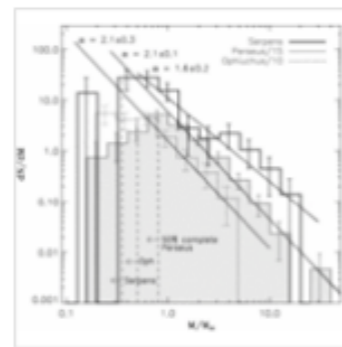
Most prominent outliers in terms of EW(10 μm) i...



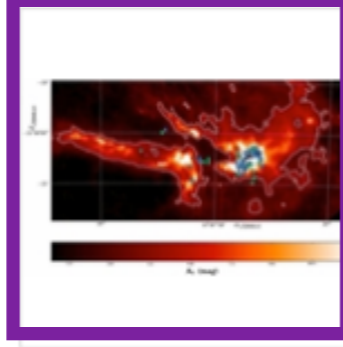
Comparison of the distribution of sizes of



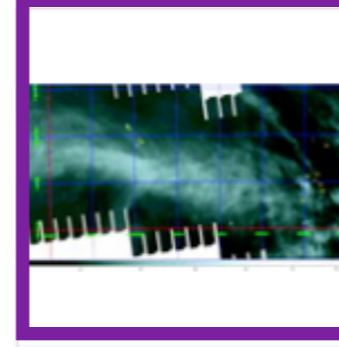
Median IRS spectra for Taurus, the **Ophiuchus** co...



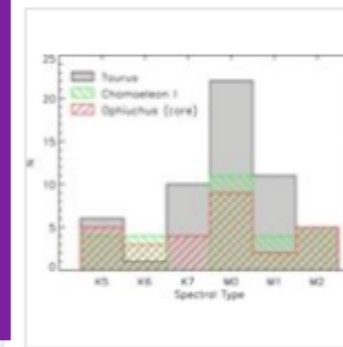
Comparison of the differential CMDs of



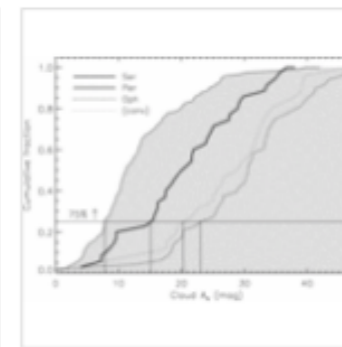
Location of the YSO (blue circles) and candidat...



High-resolution map of the **Ophiuchus** Streamer a...



Histogram for the distribution of spectral



Cumulative fraction of 1.1 mm cores as a functi...



TABLE 2
CLANS OBSERVATION SUMMARY

Target Name	Observation ID	α (J2000.0)	δ (J2000.0)	Observation Start Date	Exposure ^a (ks)
SWIRE LOCKMAN 5 (center).....	5023	10 46 00.00	+59 01 00.00	2004 Sep 12 21:30:56	67
SWIRE LOCKMAN 1	5024	10 44 46.15	+58 41 55.45	2004 Sep 16 06:53:47	65
SWIRE LOCKMAN 2	5025	10 46 39.44	+58 46 51.24	2004 Sep 17 20:30:04	70
SWIRE LOCKMAN 3	5026	10 48 32.77	+58 51 47.33	2004 Sep 18 16:17:12	69
SWIRE LOCKMAN 4	5027	10 44 06.67	+58 56 05.28	2004 Sep 20 14:40:35	67
SWIRE LOCKMAN 6	5028	10 47 53.44	+59 05 57.00	2004 Sep 23 03:36:12	71
SWIRE LOCKMAN 7	5029	10 43 27.23	+59 10 15.07	2004 Sep 24 03:43:15	71
SWIRE LOCKMAN 8	5030	10 45 20.56	+59 15 11.16	2004 Sep 25 19:47:08	66
SWIRE LOCKMAN 9	5031	10 47 13.85	+59 20 06.95	2004 Sep 26 14:47:00	65

NOTE.—Units of right ascension are hours, minutes, and seconds, and units of declination are degrees, arcminutes, and arcseconds.
^a Total good time with dead-time correction.

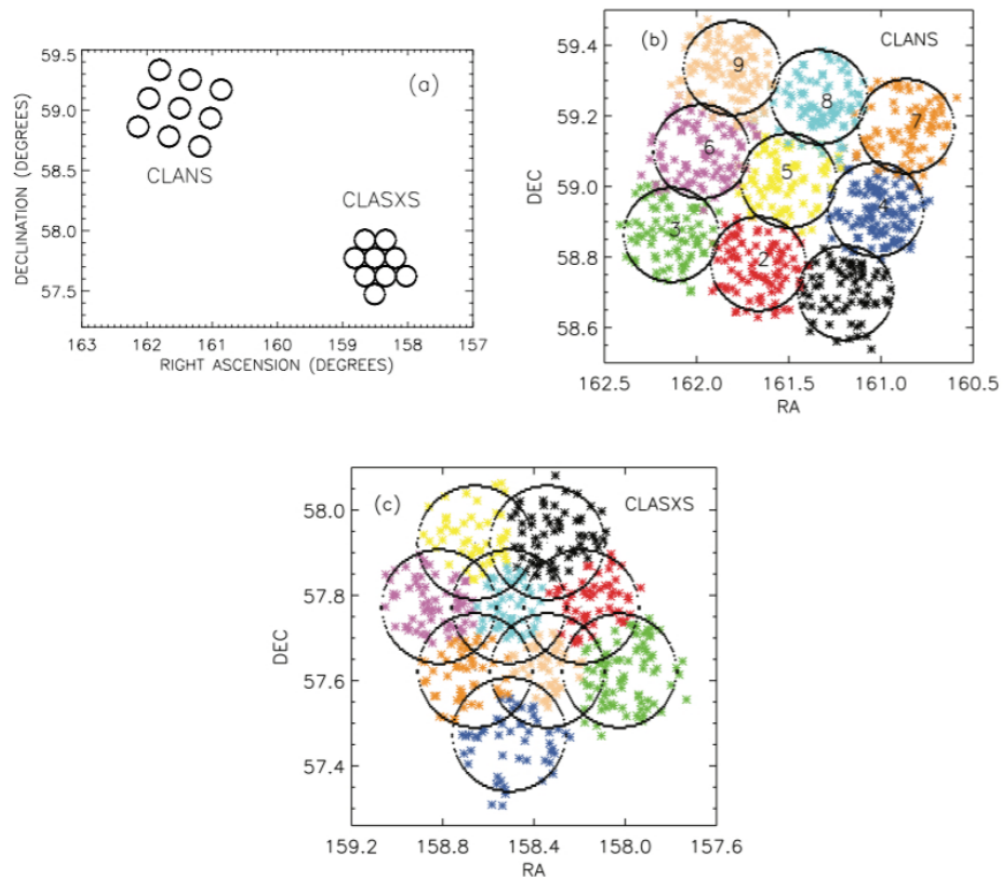


FIG. 1.—(a) Location of the CLANS and CLASXS pointings in the Lockman Hole. The circles delimit 5' from the pointing centers (the area within which the sensitivity of source detection in the ACIS-I images is approximately uniform). Location of the (b) CLANS and (c) CLASXS X-ray sources are also shown. The circles delimit 8' from the pointing centers, which is the limiting off-axis angle used in § 8 when calculating the log N –log S distributions for these fields. The numbers 1–9 in (b) correspond to the *Chandra*/SWIRE Lockman pointings listed in Table 2.

Navigation icons: mouse cursor, zoom in (+), zoom out (-), zoom reset (+), refresh (circular arrow).

How many of the figures on this journal page show a photograph of the sky, a contour map of the sky, or another type of image of the sky *with the axes labeled*?

- 0
- 1
- 2
- 3
- 4
- 5+

Need some help with this task?

Back

Done



Show the project tutorial

[demo]

1610



SIDEREUS NUNCIUS

On the third, at the seventh hour, the sequence. The eastern one was 1 minute, the closest western one 2 minutes; and the

East * ○ * * West

to minutes removed from this one. They were absolutely on the same straight line and of equal magnitude.

On the fourth, at the second hour, there were four stars around Jupiter, two to the east and two to the west, and arranged precisely

East * ○ * * West

on a straight line, as in the adjoining figure. The easternmost was distant 3 minutes from the next one, while this one was 40 seconds from Jupiter; Jupiter was 4 minutes from the nearest western one, and this one 6 minutes from the westernmost one. Their magnitudes were nearly equal; the one closest to Jupiter appeared a little smaller than the rest. But at the seventh hour the eastern star was 30 seconds apart. Jupiter was 2 minutes from the

East ** ○ **

one, while he was 4 minutes from the next western one was 3 minutes from the westernmost one. They and extended on the same straight line along the ecliptic.

On the fifth, the sky was cloudy.

On the sixth, only two stars appeared flanking Jupiter

East * ○ *

in the adjoining figure. The eastern one was 2 minutes from the next western one 3 minutes from Jupiter. They were on the same straight line with Jupiter and equal in magnitude.

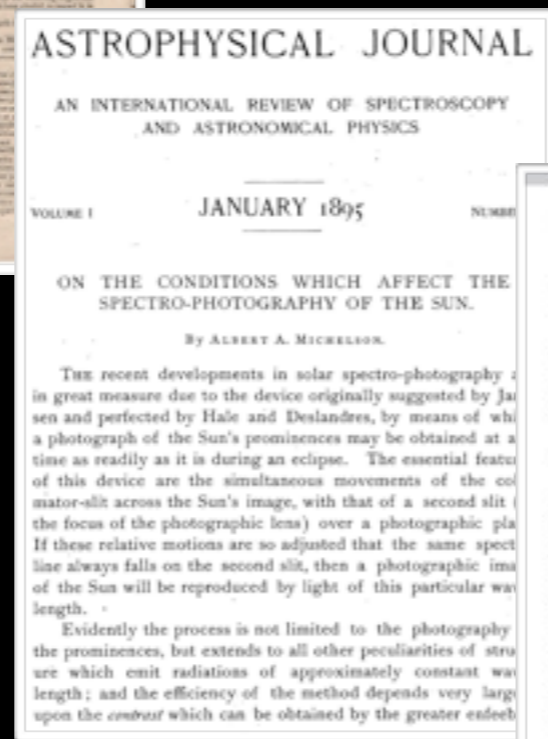
On the seventh, two stars stood near Jupiter, but not on the same straight line with Jupiter and equal in magnitude.

1665

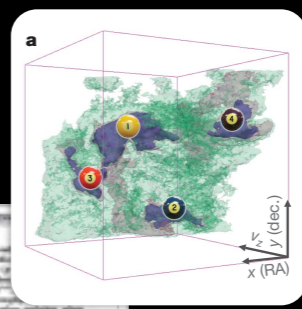


SCHOLARLY COMMUNICATION

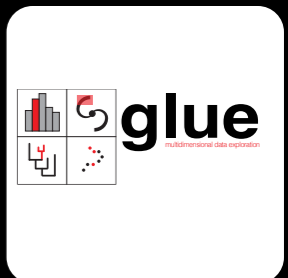
1895



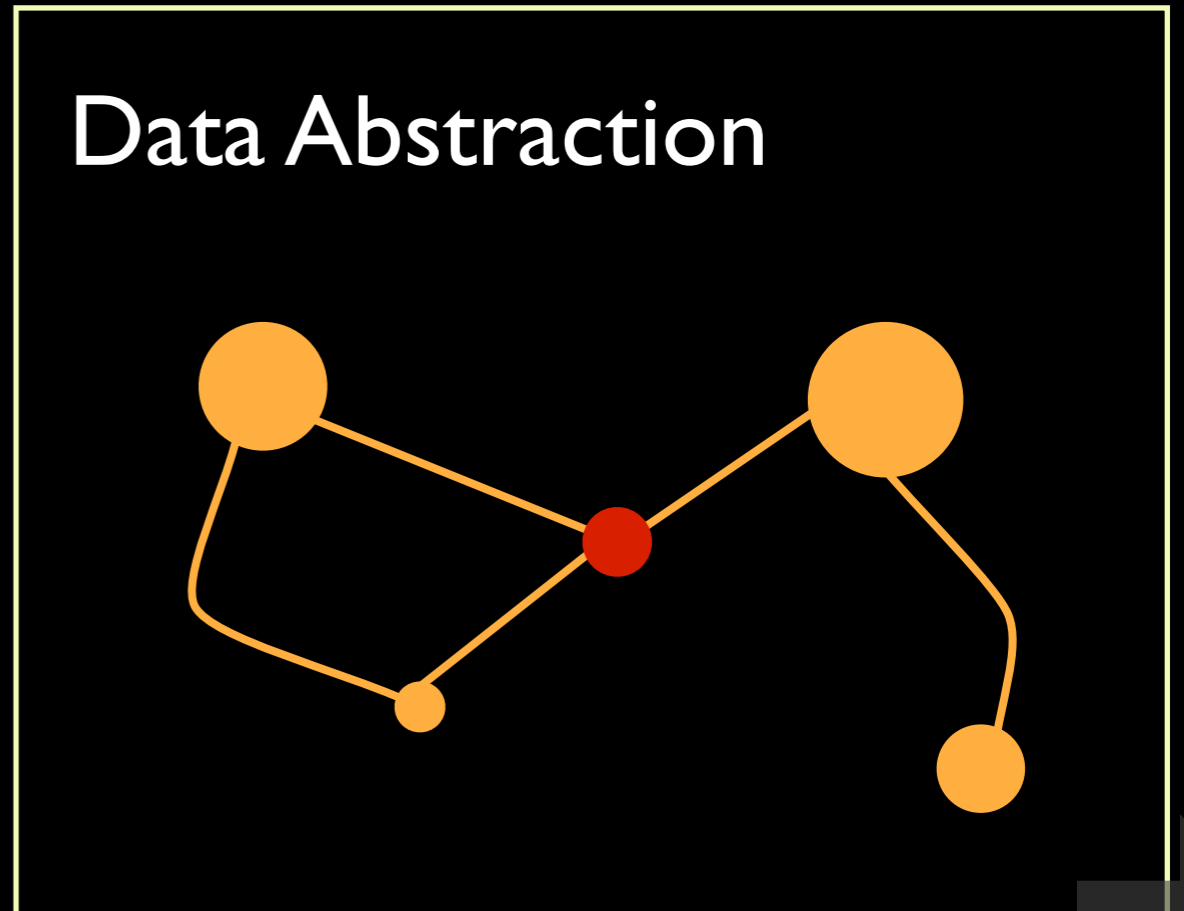
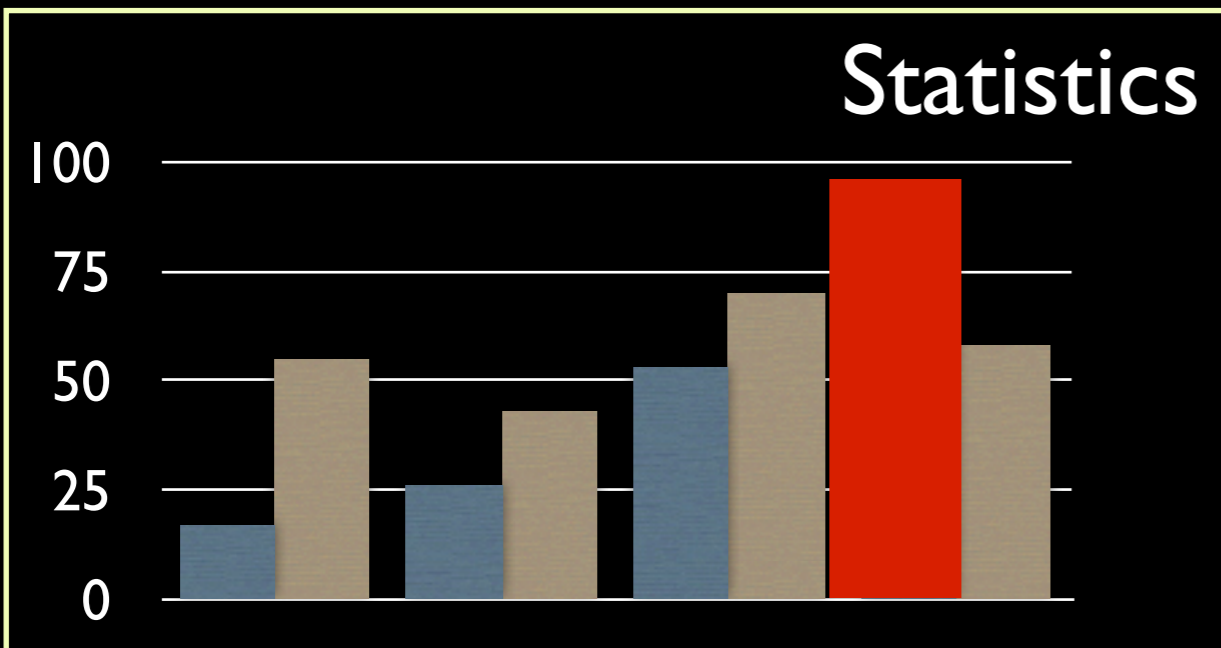
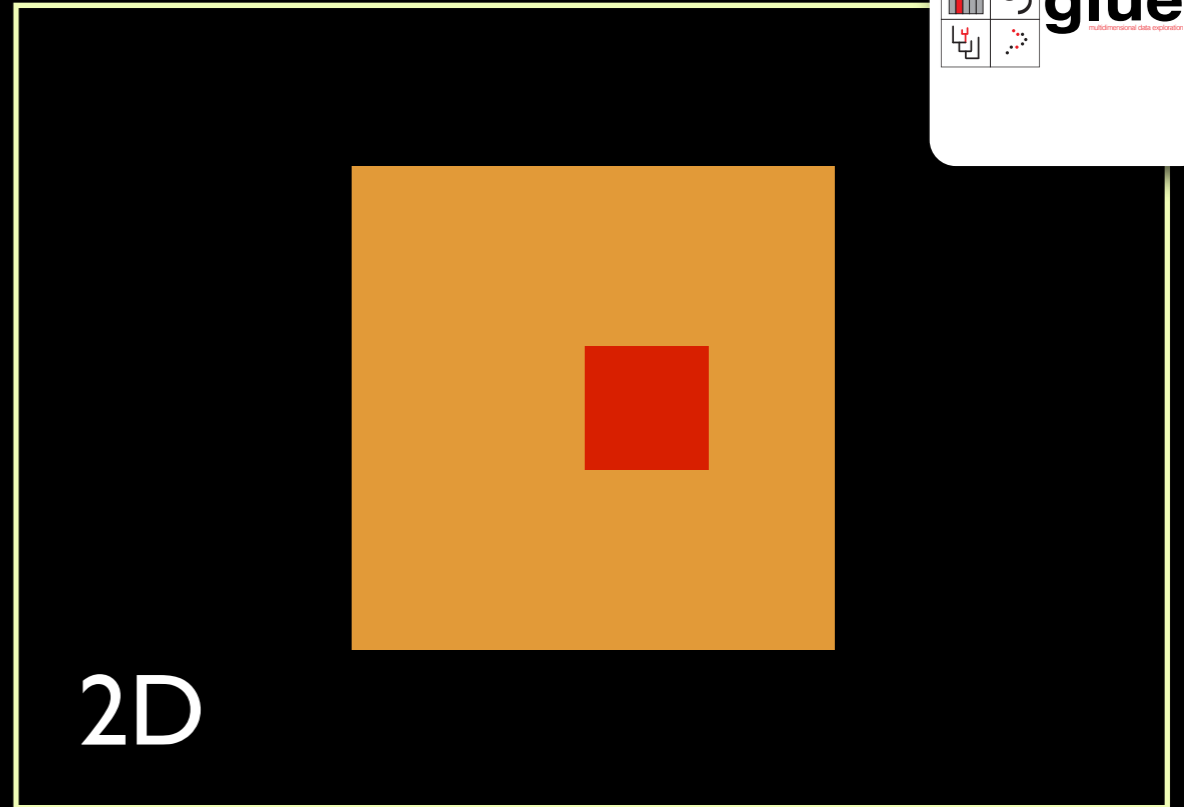
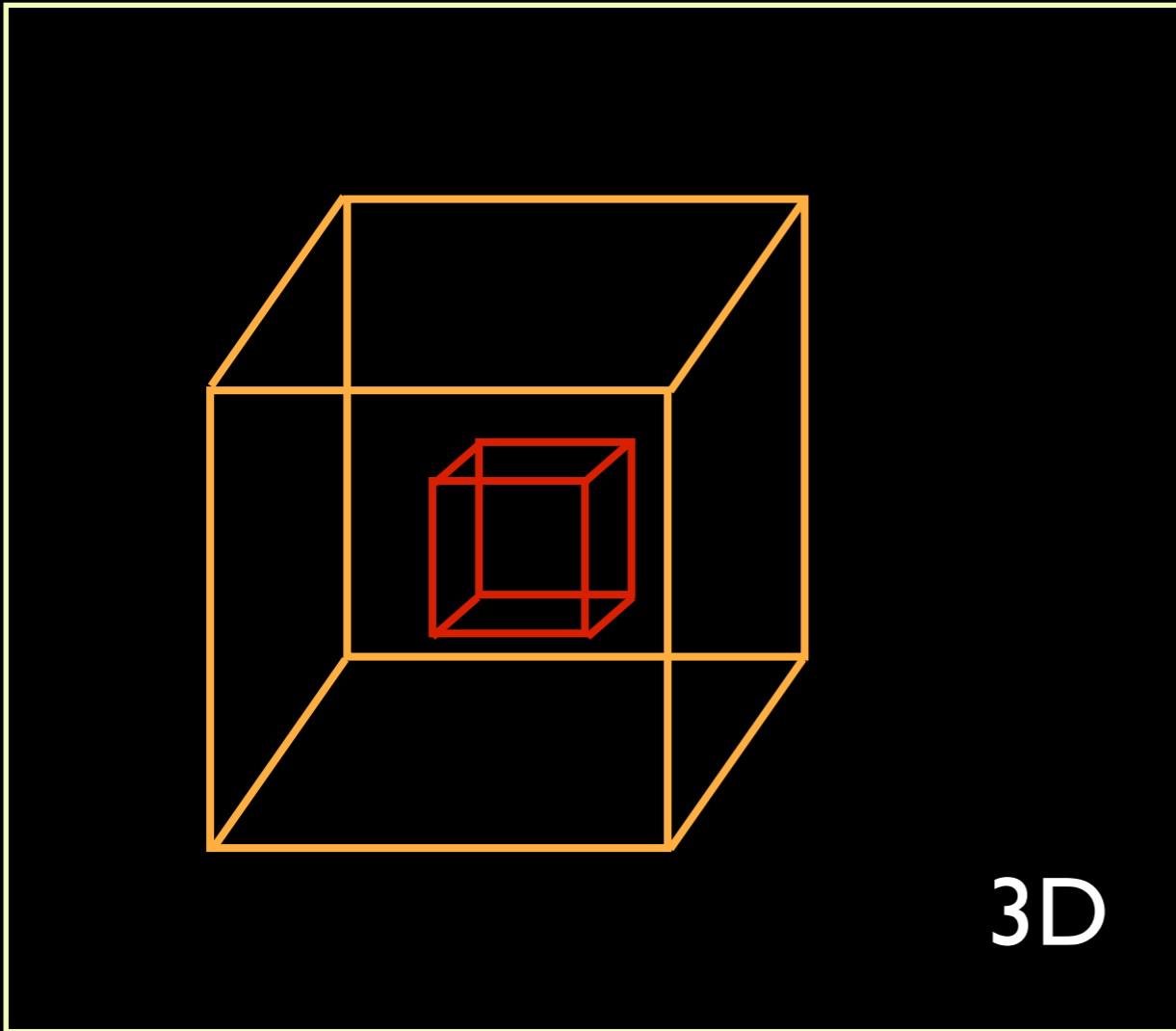
2009



2015



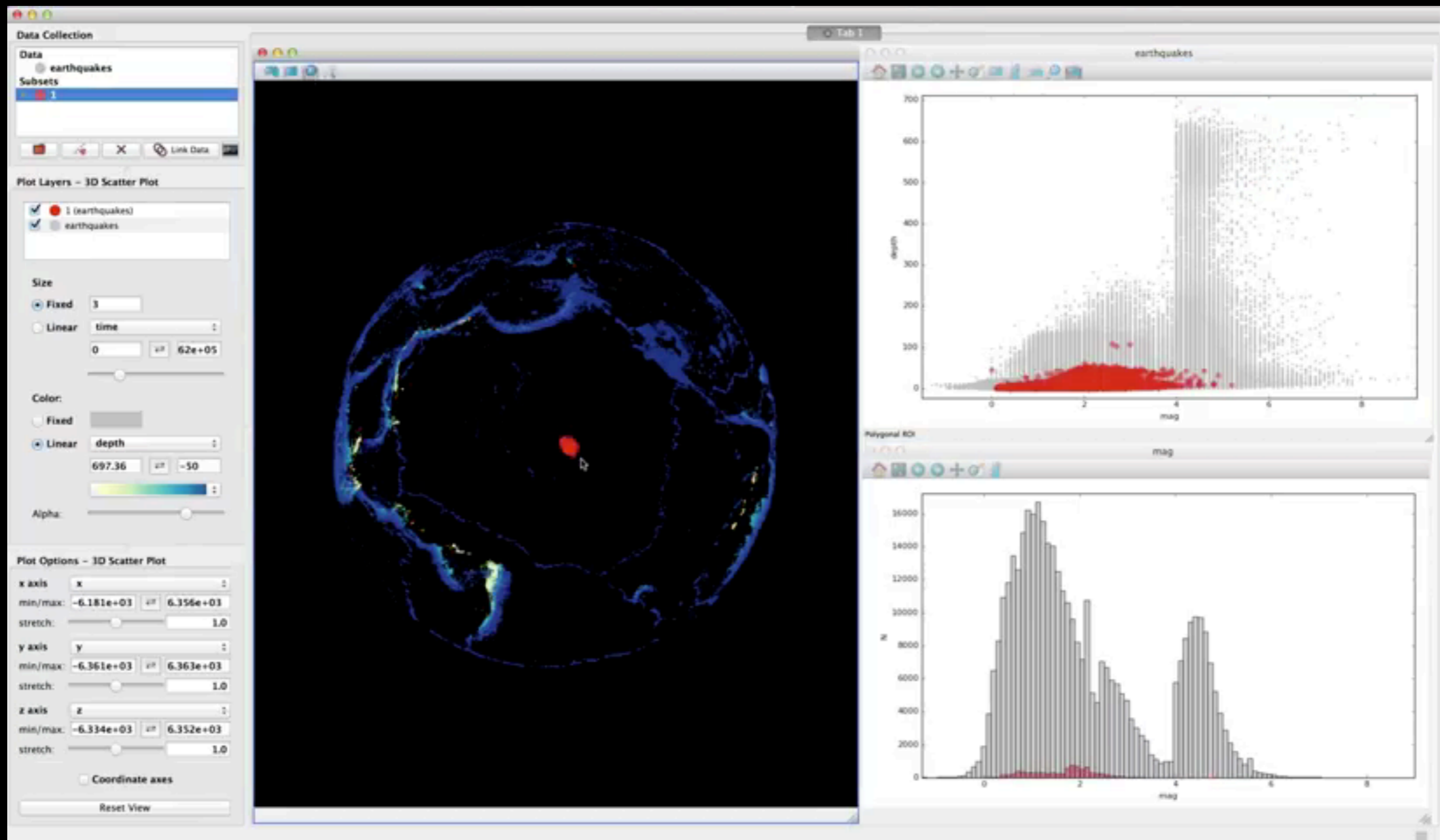
LINKED VIEWS OF HIGH-DIMENSIONAL DATA



figure, by M. Borkin, reproduced from [Goodman 2012](#), "Principles of High-Dimensional Data Visualization in Astronomy"

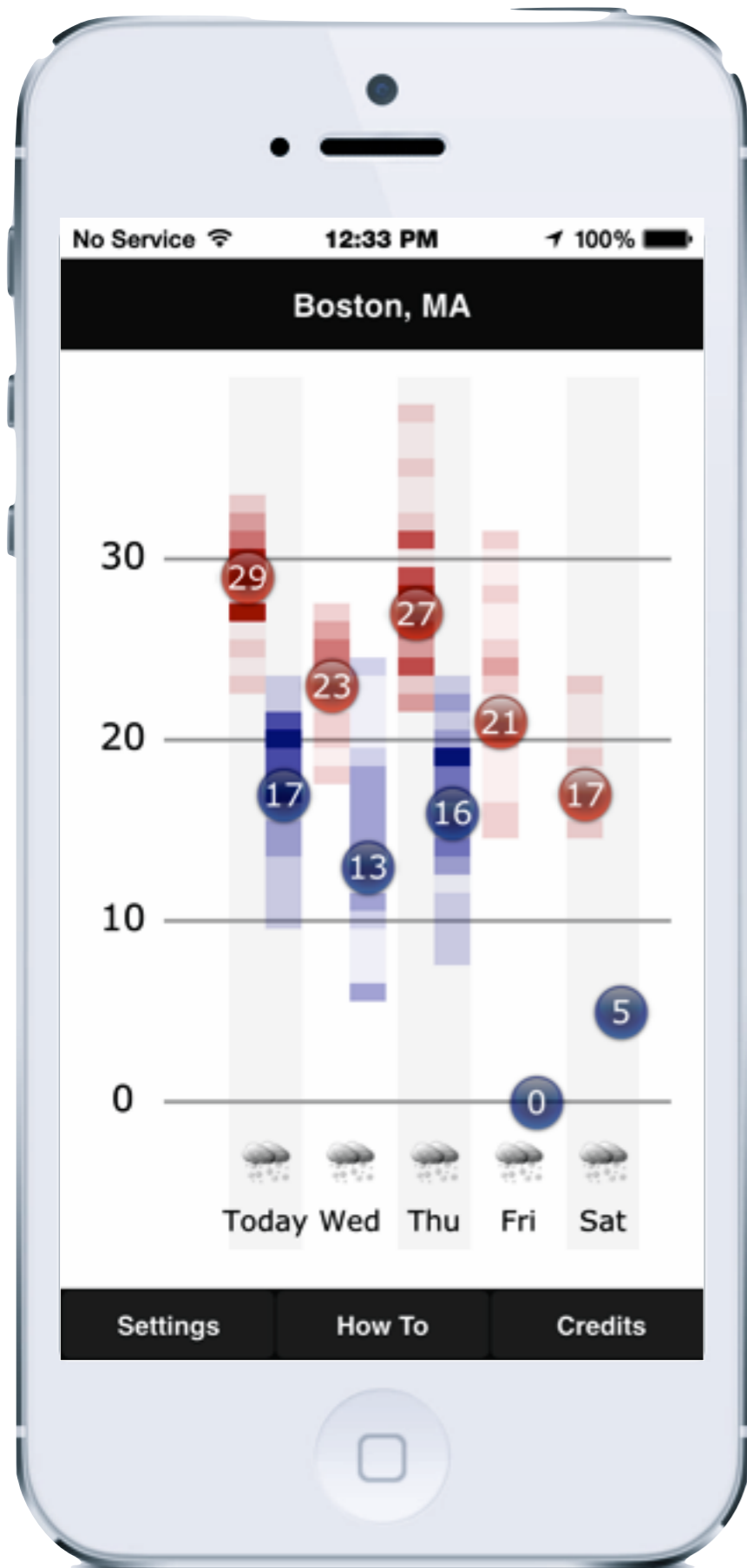
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GLUE



*video by Tom Robitaille, lead glue developer
glue created by: C. Beaumont, M. Borkin, P. Qian, T. Robitaille, and A. Goodman, PI*

What am I doing tomorrow?



PREDICTIONX



App Store > Weather > Harvard University

Take A Sweater

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Rating: 4+
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NOTE: Take-A-Sweater currently only has data for Boston, MA. This will be changing with the next release.

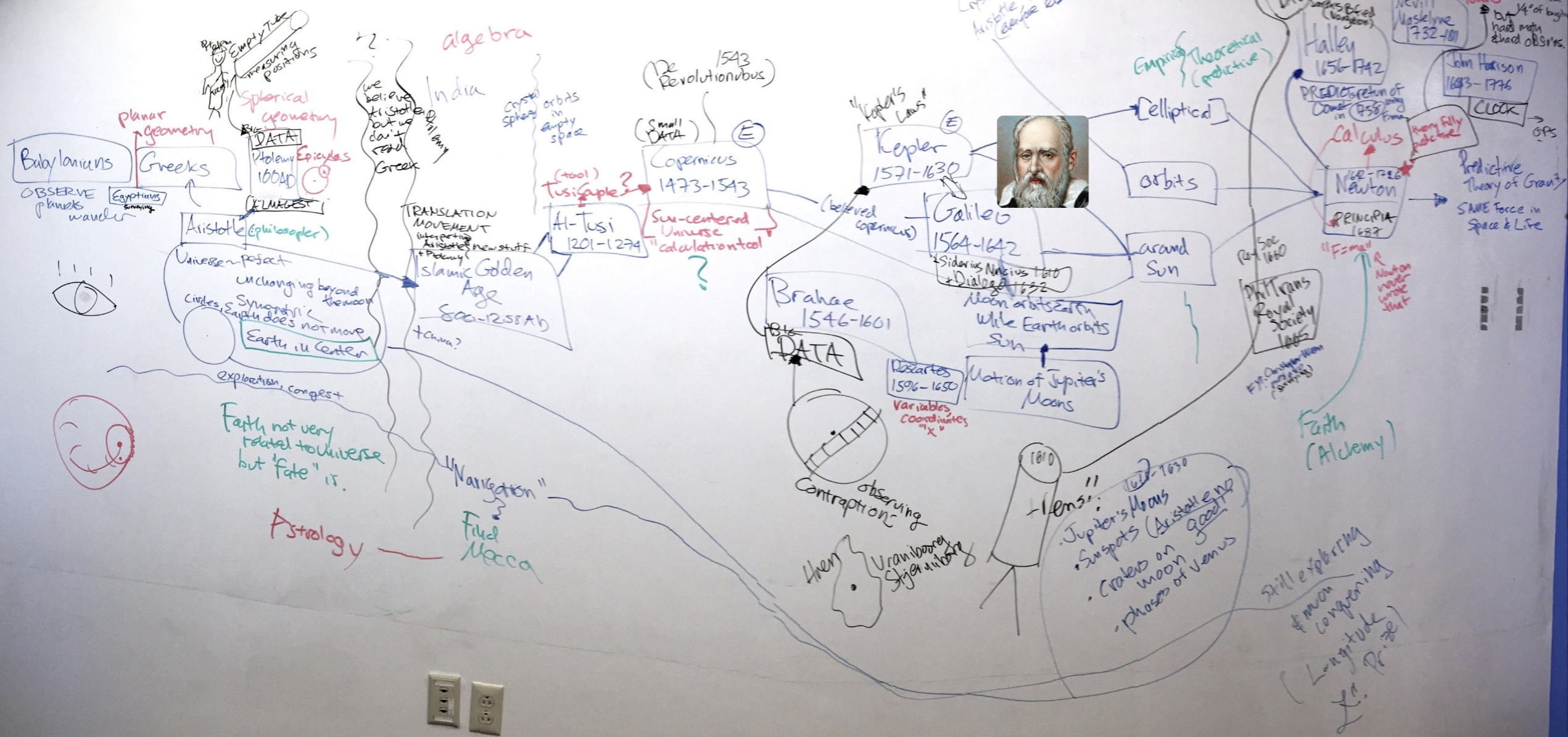
This App was created in 2012, for use in the Harvard University General Education course "The Art of Numbers," taught by Prof. Alyssa Goodman. The code was written by Bill Barthelmy of Harvard's Academic Technology Group. Historical data were kindly provided by ForecastWatch, a product of Intellovations, LLC. Current five-day weather forecast data are provided by NOAA....

takesweater.com, and "TakeASweater" in the Apple App Store



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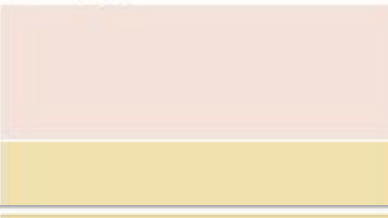
Solar System = Universe
Earth = center of Universe





PREDICTIONX

Geographic Region Picker



400

500

600

700

800

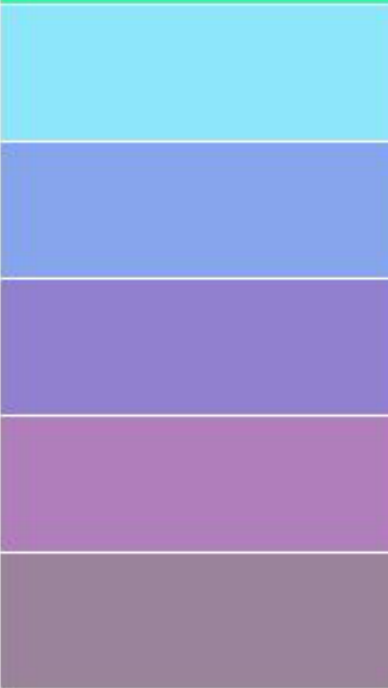
900

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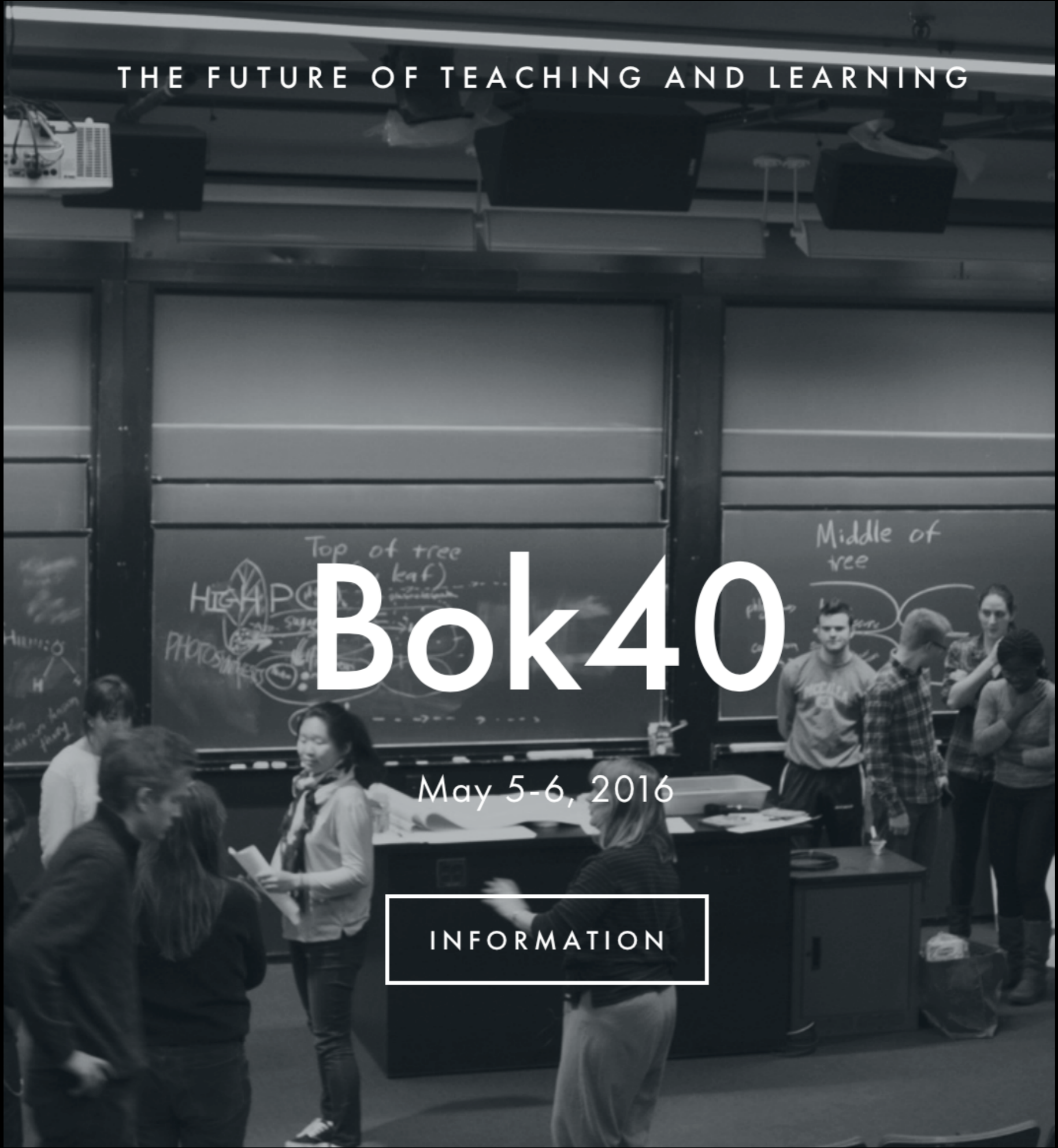
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THE FUTURE OF TEACHING AND LEARNING

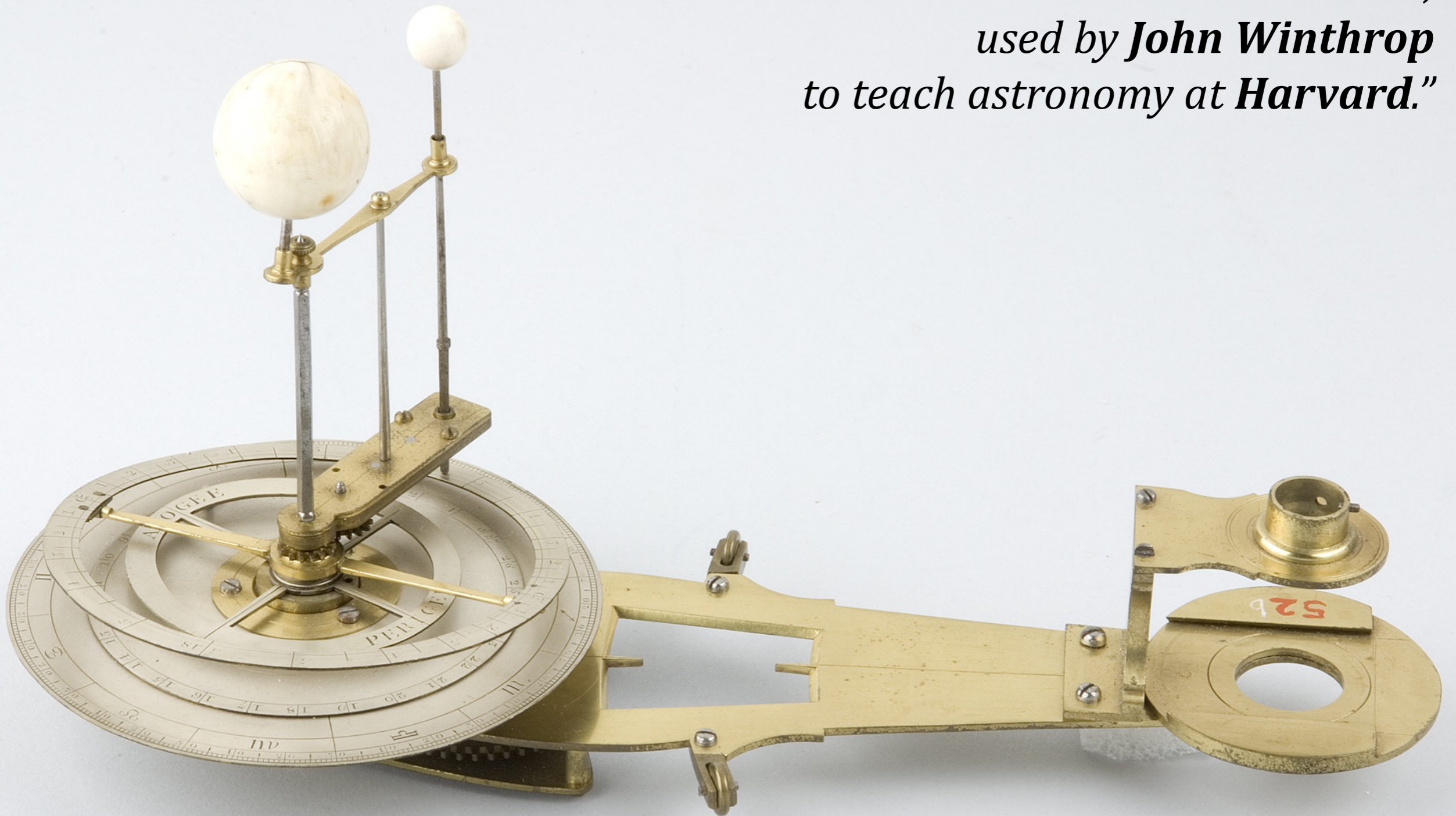
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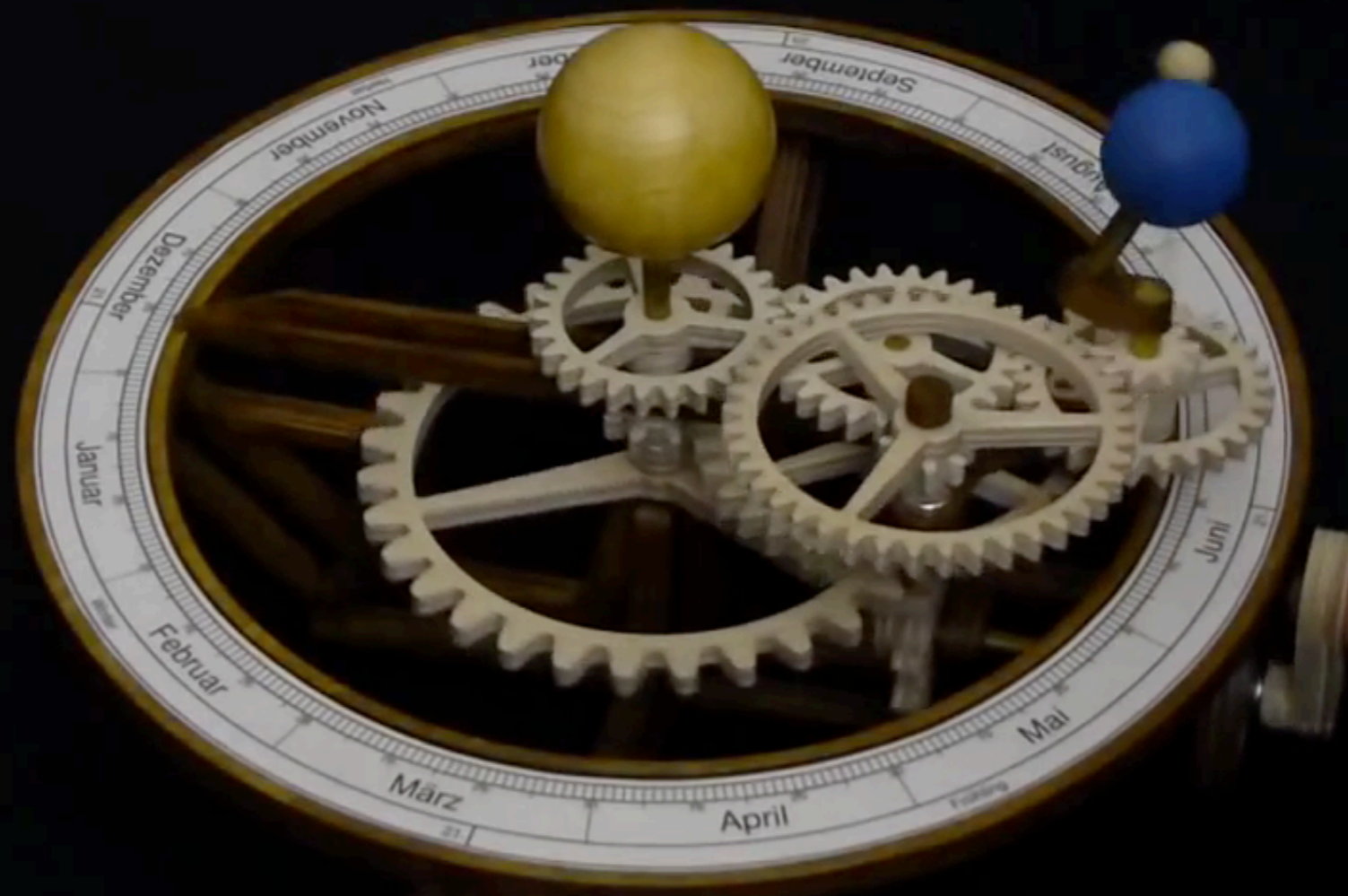
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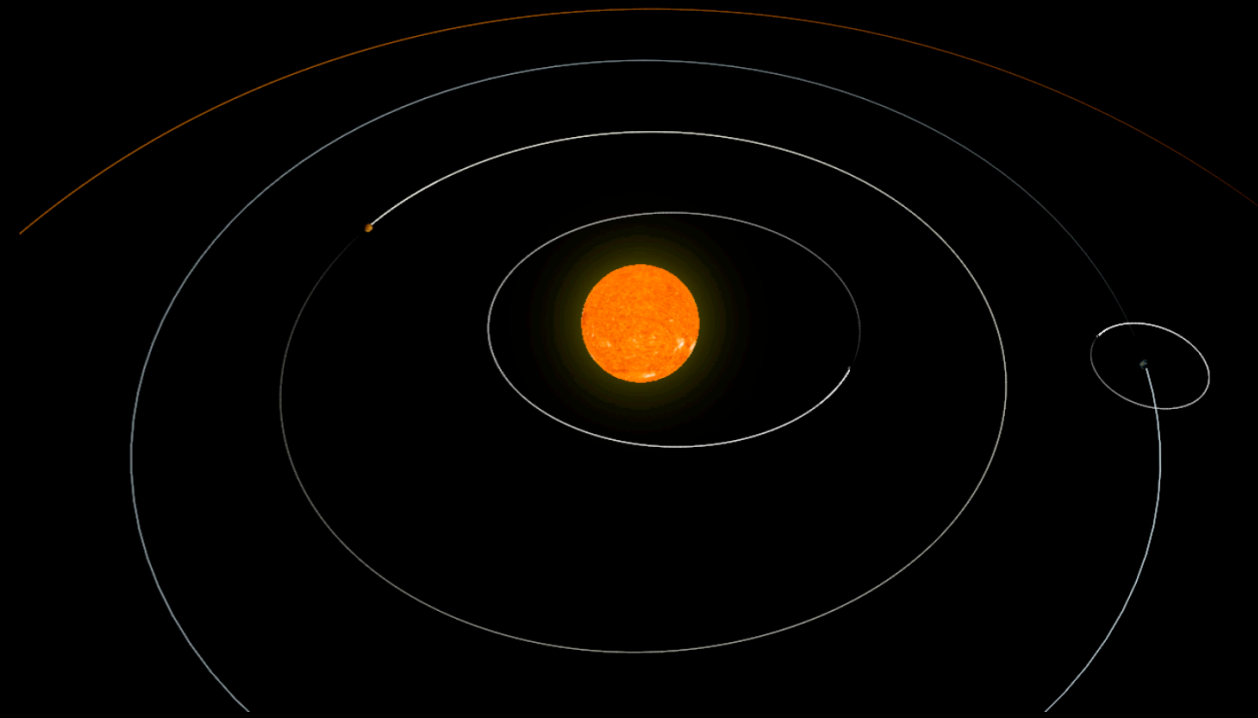
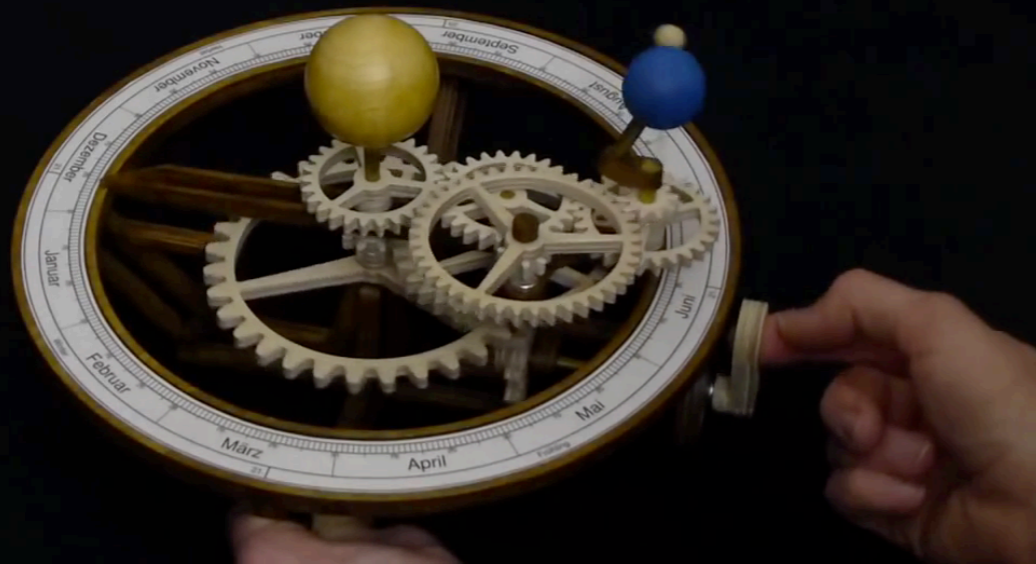


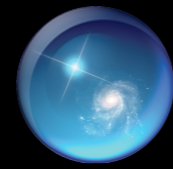
*“A tellurion made in 1766,
used by John Winthrop
to teach astronomy at Harvard.”*



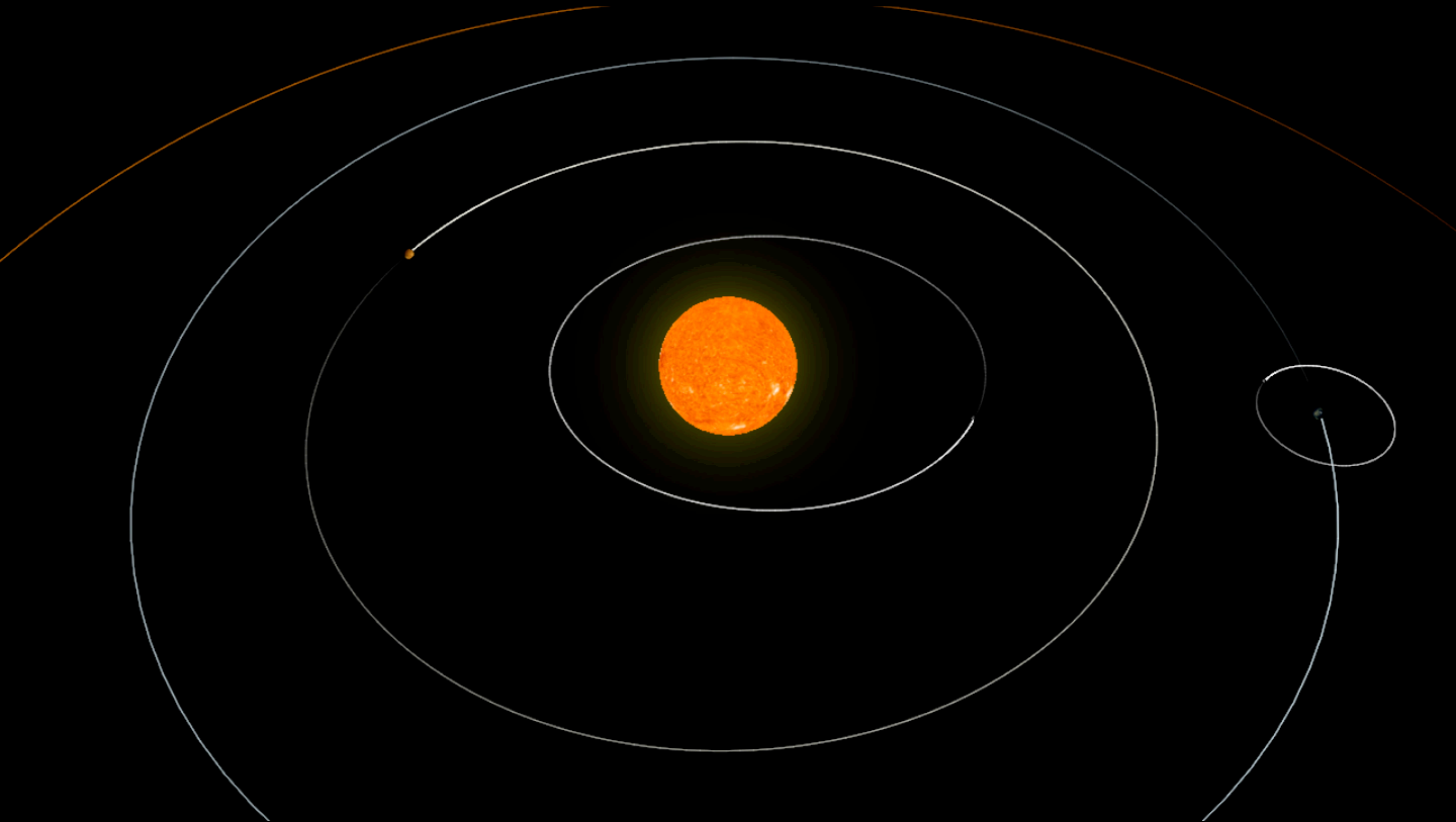


Tellurion, Holz Mechanik

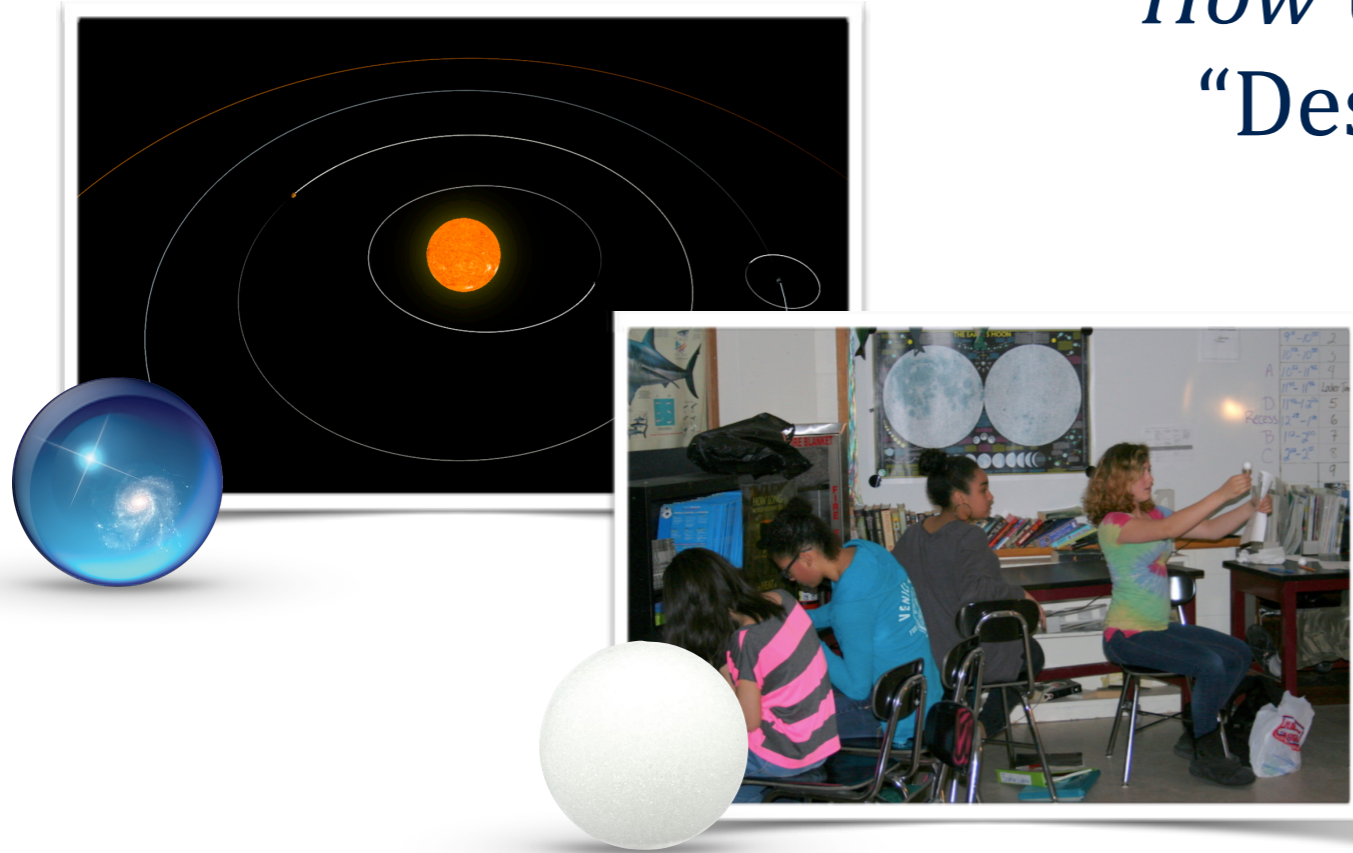




WorldWide Telescope



How to Mix “Design”



Sample Student Drawing



Pre



Post

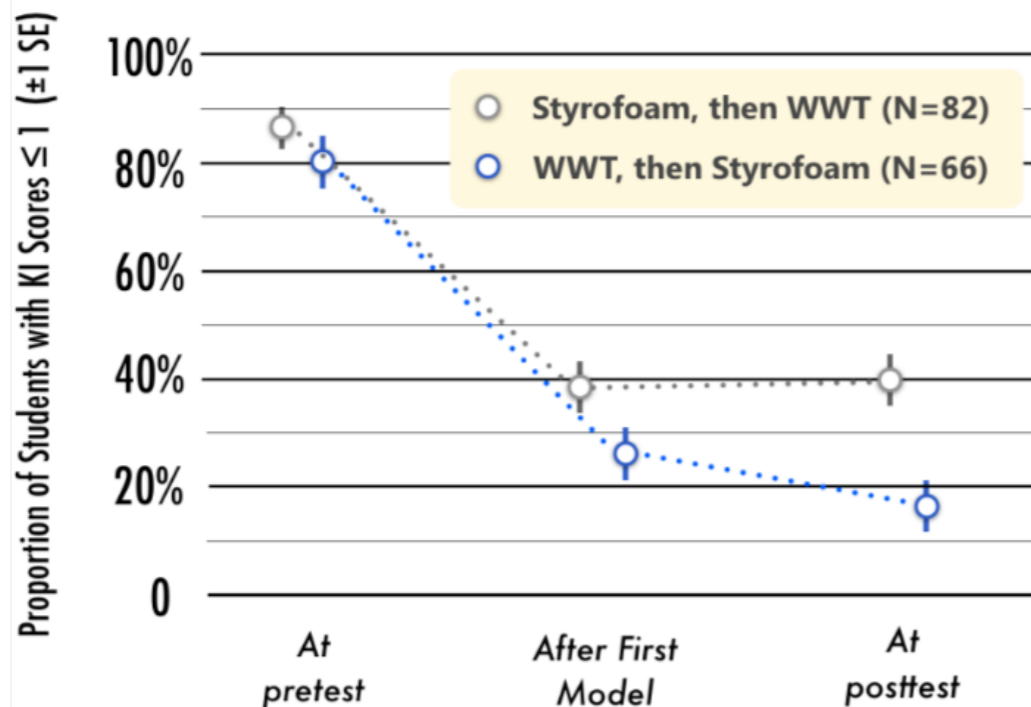
Sample Student Response

We see a half Moon on this day because:

(Pre) “The earth is blocking part of the moon from the sun and it casts a shadow on the moon, making half of it dark.”

(Post) “We are seeing 1/2 of the lit up part of the moon. If a line is drawn on the moon to show our perspective, we see half of the light of the moon.”

Students whose KI Scores Indicate Misconceptions or No Scientifically Valid Responses (A13 & A14)



Model Preference	Proportion of Students with that Preference (All Phase 2, N=294)
Models helped equally	37% ± 3%
Computer helped more	40% ± 3%
Styrofoam helped more	13% ± 2%

How to Mix "Design"

online learning

peer instruction

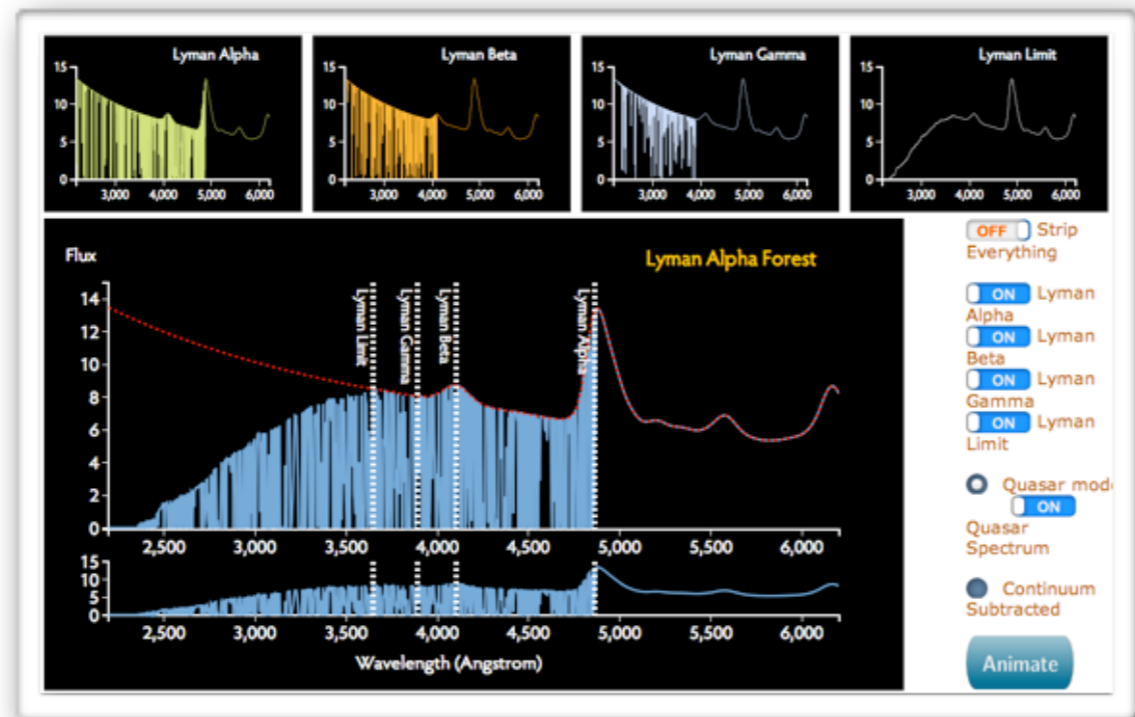


Stephen



Yuan-Sen Ting

Interstellar Absorption and the Lyman Alpha Forest



JavaScript JavaScript

https://www.cfa.harvard.edu/~yuan-sen.ting/lyman_alpha.html

JavaScript JavaScript

<http://portillo.ca/nebula/>



virtual environments

